

# Subaru High-Contrast Imaging Are Critical for Technical Demonstration and Scientific Studies of the Roman Coronagraph

Proposers (name/affiliation): M. Kuzuhara (ABC/NAOJ), S. Miyazaki (ISAS/JAXA), T. Mizuki (ABC/NAOJ), T. Uyama (CSU), N. Murakami (ABC/NAOJ), S. Wolf, J. Hom (U. of Arizona), J. Girard (STScI), J. Wang (Northwestern U.), V. Bailey (JPL), M. Millar-Blanchaer (U. of California), Roman CPP Team, and SCExAO team

Subaru Instrument: SCExAO/CHARIS, REACH, VAMPIRES (possibly, exo-NINJA)

Number of nights (hours): 30 nights

Condition of nights (moon phase, airmass, seeing): no restriction for moon phase, airmass  $< 2.0$ , seeing  $< 1.0''$

Time critical (year, season, date, time): in 2027 (= Phase 1: 25 nights) + in 2028–29 (5 nights = Phase 2)

Relevant CCS/other Roman program: Coronagraph Instrument (high contrast imaging)

Category (exoplanet, galaxies, large scale structure, solar system, stellar physics, stellar population/ISM, super massive blackhole/AGN, IGM/CGM) : exoplanet

Key words: high contrast direct imaging, planet, brown dwarf, atmosphere, spectroscopy

# Goals and Targets of Roman High Contrast Imaging

- **Roman-Coronagraph: technical demonstration for a future mission that is for imaging an Earth twin (i.e., HWO), but available for sciences after the first observation phase.**
- First demonstration goal: achieve a contrast of  $10^{-7}$  -  $10^{-8}$  at 580 nm over a dark-hole area ( $= 6-9 \lambda/D$ ) with a total integration of few hours for a very bright star ( $V < 5-6$ )
  - It is essential to test the performance by detecting an actual companion with a contrast of  $10^{-7}$  -  $10^{-8}$ , rather than an evaluation based on noise level.
  - **The detection of a companion having the above contrast level enables scientific studies simultaneously as the demonstration.**
  - The **samples must be located within  $5^\circ$  in Ecliptic longitude** from bright reference stars ( $V < 3$ ; diameter  $< 2$  mas), which are observed for modeling a dark hole but **sparse over the sky**
  - **Few known samples** meet the above requirements.
  - The Subaru team is now conducting a survey to find substellar companions around bright stars, potentially providing the samples suitable for the Roman (below, **this survey = the ongoing survey**)

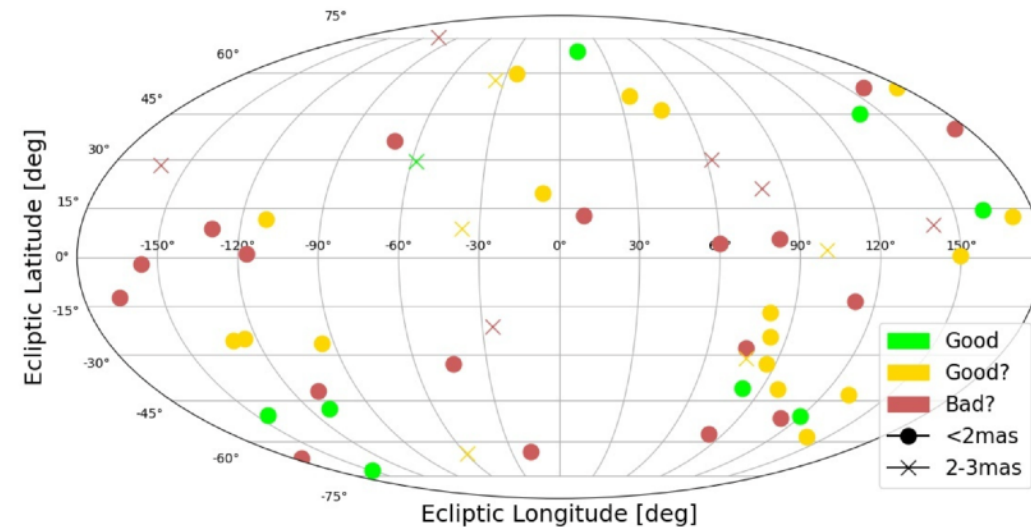
# Subaru SCExAO Survey for Roman Targets: Phase 1

- Only our ongoing survey (see 2nd page) may not be able to find enough samples, due to the restrictions from e.g.
  - the visibility of targets observable in time-window allocated to the Roman coronagraph project
  - Small separation angles between references and science targets

We propose a Roman-Subaru synergetic program to increase the samples

- 25 nights are requested to discover 5 substellar companions that are suitable for the Roman coronagraph
- Adopt the target-selection policy optimized for the Roman coronagraph: all the targets (N~60) will be selected from the young stars that are appropriately close ( $< 5^\circ$ ) to reference stars
- We expect leveraging Gaia DR4 data boost discovery prospects by several-fold

Map of selected dark-hole references (credit: J. Hom)

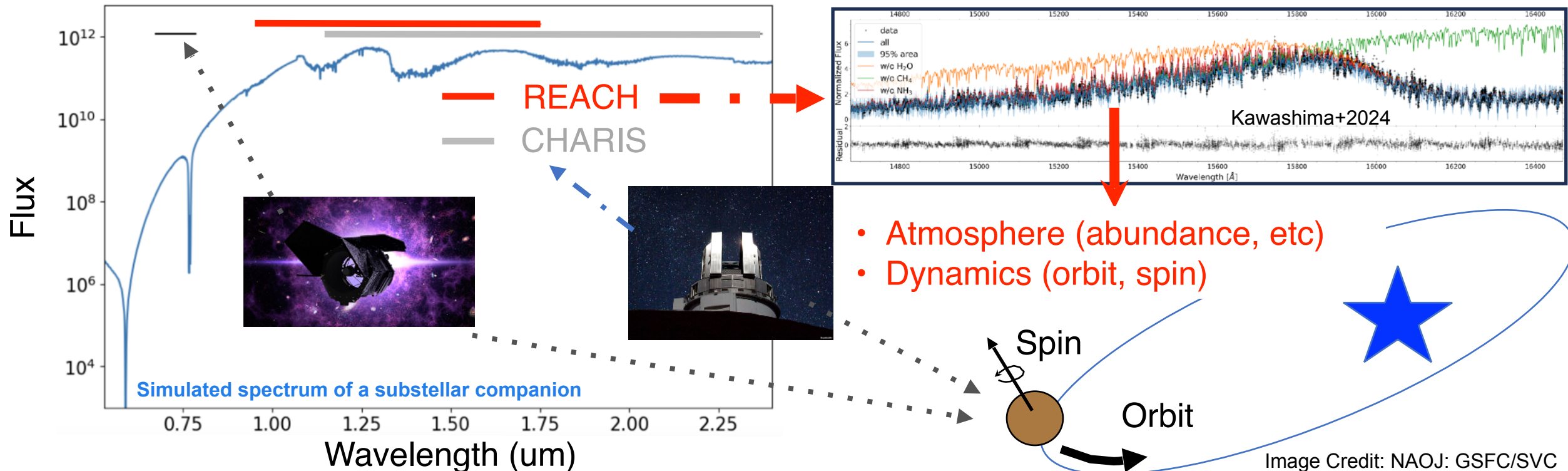


We are vetting candidates of reference star now miro

# Follow-Up Characterizations: Phase 2

- Request additional 5 nights for in-depth characterization of ~5 best samples from our surveys, which we also propose to observe with Roman (note 1st phase + Roman already enables some characterizations)
- Reveal atmosphere, orbit, and obliquity of substellar companion by comprehensively analyzing all the imaging and spectroscopy data

## Spectroscopic and imaging characterization with Subaru and Roman



# Summary

- The targets for the Roman coronagraph should be very close ( $\Delta < 5^\circ$  in Ecliptic longitude) to dark-hole reference stars, which are sparse over the sky
- It is important to create the sufficient list of substellar companions suitable for demonstrating the Roman's high-contrast performance, to avoid a risk that there will be no targets available in the time allocated for the coronagraph project of Roman
- Discover  $\sim 5$  substellar companions using Subaru and image them at visual wavelengths using Roman, fulfilling one of the Roman's missions and enabling the first visible-to-infrared characterization of a substellar-mass companion (e.g., constraint on cloud, composition)
- Characterize companions discovered by Subaru and followed up by Roman, using the unique capability of the Subaru Telescope for high-contrast spectroscopy

**Significance of Synergy:** Subaru observations are crucial not only for achieving a key goal of Roman's coronagraph project but also for advancing scientific studies via Roman's high-contrast imaging and for fully characterizing substellar companions observed by Roman.