#### Probing AGN Demographics with Roman+Subaru

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Subaru Instrument: PFS, HSC, NINJA

Number of nights (hours): 50 nights

Condition of nights: dark/grey

Time critical (year, season, date, time): N/A

Relevant CCS/other Roman program: SPQR (WFS, PI: J. Rhoads)

Category: supermassive black holes

# High-redshift Quasar

Roman: HL Wide Area Survey (DEC > -30 deg) + Time Domain Survey

Subaru: PFS/NINJA + HSC

◆ Science Goals: early growth history of supermassive black holes (SMBH)

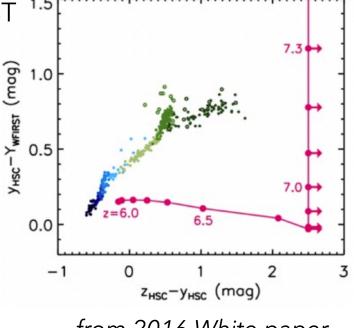
(1) Discovery of z>7 quasars (~10 Gpc<sup>-3</sup>, or several per 50 deg<sup>2</sup> at Y<25)

- PFS confirmation of quasar candidates selected from Roman + LSST

- HSC z/y imaging (26 mag;  $5\sigma$ ) in Deep survey area

- PFS/NINJA also enable BH mass estimate with CIV and MgII

- (2) Large-scale environments
  - PFS identification of surrounding galaxies to characterize local overdensity



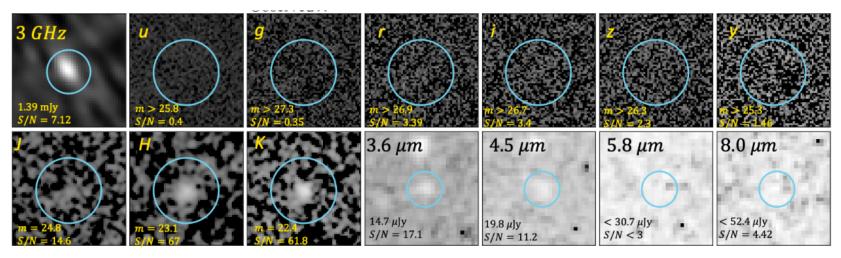
from 2016 White paper (Y. Matsuoka)

### Dust-reddened AGN

Roman: HL Wide Area Survey + Time Domain Survey

Subaru: HSC

- ◆ Science Goals: Role of dust obscuration for AGN, SMBH demographics
- (1) NIR counterparts of optical-dark AGN
  - 8 HSC-dark VLASS 3GHz sources in u2k catalog (12-17 deg<sup>2</sup>)
    - → Statistical studies possible with 50-100 deg² coverage with HSC-D/UD depth
    - → Roman's deep NIR imaging helps identify dusty radio AGN or very high-z AGN



## Quasar Host

Roman: HL Wide Area Survey + Time Domain Survey

Subaru: PFS/NINJA

- ◆ Science Goals: Co-evolution of galaxy and SMBH
  - \* Roman's image depth and quality are better suitable for imaging decomposition studies of quasar hosts than Euclid
  - \*Roman's wide coverage enables statistics of quasar hosts at low-mass end than JWST
- (1) Stellar velocity measurements at z<2.1
  - Targets from Roman + LSST
  - Stellar mass from Roman
  - redshift,  $M_{BH}$  and  $\sigma_*$  from PFS

#### Follow-up of Prism/Grism & Narrowband sources

Roman: HL Wide Area Survey + Time Domain Survey

Subaru: PFS/NINJA

◆ Science Goals: Characterization of broad-line sources from Roman's spectroscopic data & HSC narrowband imaging survey

- (1) AGN candidates from PRISM/GRISM
  - Confirmation of broad-line AGN candidates
  - M<sub>BH</sub> through Balmer & metal line profiles
- (2) AGN candidates from narrowband imaging
  - AGN would dominate in luminous line emitters
- → Contribution of broadline AGN to Ha / Lya luminosity function

## Summary

#### Significance of Synergy:

- Subaru's wide-field survey power and the unique capability of HSC+PFS+NINJA enable discovery of rare AGN populations (high-z quasar, dusty AGN)
- With Roman's deep sensitivity and high imaging resolution,
  one can study quasar hosts better than Euclid