

# SPICA Mid-Infrared Camera and Spectrometer

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“spica\_mir” team

# Instrument Overview

5 -- 38 $\mu$ m(--50 $\mu$ m) Camera and Spectrometer

- Wide Field Camera

- 5 arcminutes square FOV x 2,  $\lambda\lambda$  5--25 and 20--38 $\mu$ m

- High Resolution Spectrograph

- $R \sim 30,000$   $\lambda\lambda$  4--8  $\mu$ m and 12--18 $\mu$ m

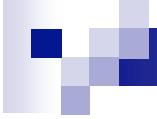
- Mid Resolution Spectrograph

- IFU by image slicer

- $R \sim 1500$ --700  $\lambda\lambda$  (10--20)+(20--36) $\mu$ m at once

- Low Resolution Spectrograph

- $R \sim 50$ --100  $\lambda\lambda$  5-26 $\mu$ m and (20-38 or 25-48) $\mu$ m



# Membership

Hirokazu Kataza  
Takehiko Wada  
Itsuki Sakon  
Naoto Kobayashi  
Masayuki Akiyama  
Eiichi Egami  
Fumi Egusa  
Keigo Enya  
Naofumi Fujishiro  
Tomotsugu GOTO  
Poshak Gandhi  
Yashuhiro Hirahara  
Mitsuhiko Honda  
Takashi Ichikawa  
Yuji Ikeda  
Myungshin Im  
Masatoshi Imanishi  
Daisuke Ishihara  
Yoshifusa Ita  
Hideyuki Izumiura

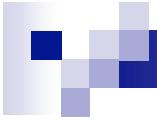
Masakazu KOBAYASHI  
Hidehiro Kaneda  
Mitsunobu Kawada  
Hideyo Kawakita  
Tadayuki Kodama  
Takayuki Kotani  
Yusei Koyama  
Takashi Kozasa  
Hideo Matsuhara  
Toshio Matsumoto  
Shuji Matsuura  
Mikako Matsuura  
Takashi Miyata  
Hiroshi Murakami  
Tohru Nagao  
Hirohisa Nagata  
Takao Nakagawa  
Makitsubo Nobuhiro  
Takaya Nozawa  
Yoko Okada

Yoshiko Okamoto  
Takashi Onaka  
Takahumi Ootsubo  
Shinki Oyabu  
Yoichi Oyama  
Chris Pearson  
Tomoki SAITO  
Shigeyuki Sako  
Mai Shirahata  
Toyoaki Suzuki  
Toshinobu Takagi  
Michihiro Takami  
Tsutomu T. Takeuchi  
Hitoshi Tokoro  
Koji Tsumura  
Kentaroh Watanabe  
Toru Yamada  
Takuya Yamashita



## Scientific objectives

- Resolution of Birth and Evolution of Galaxies
- Thorough Understanding of Planetary System Formation
- The Transmigration of Dust in the Universe



## --- Today's topics ---

We will show

- Specifications of the instrument
- Current design
- Status of developments

and also talk about

- Feasibility of the specifications
- Priority in the development process.

Please discuss

- What is important in the future?
- Modifications of the specifications.

## Spec.: Wide-Field-Camera (WFC)

- FOV: 5' x 5' x 2 field WFC-S and WFC-L
- Diffraction limited image
- Zodiacal light limit noise
- WFC-S
  - 5 -- 25 $\mu$ m
  - Si:As 2048x2048 pixels 0."146 fov/pix
- WFC-L
  - 20 -- 38 $\mu$ m
  - Si:Sb 1024x1024 pixels 0."293 fov/pix

# Spec.: High Resolution Spectrograph (HRS)

- Immersion Grating based High Resolution

- HRS-L

- Si:As 2048 x 2048 detector
  - 12 -- 18  $\mu\text{m}$  R~20,000 -- 30,000
  - 6".0 x 1".2 long slit
  - 0".48 /pix

- HRS-S

- Si:As 2048 x 2048 detector
  - 4 -- 8 $\mu\text{m}$  R~30,000
  - 3".5 x 0".72 long slit
  - 0".288 /pix

# Spec.: Mid Resolution Spectrograph (MRS)

- Integral Field Unit by Image slicer
- 2-octave wavelength coverage by dichroic filter
- MRS-L
  - Si:Sb 1024x1024 detector
  - 20 -- 36  $\mu\text{m}$  R~700
  - 12".0 x 2".5 slitlet x 5 = 12" x 12" FOV
  - 0".485 /pix
- MRS-S
  - Si:As 2048x2048 detector
  - 10 -- 20 $\mu\text{m}$  R~1500
  - 12" x 1".2 slitlet x 5 = 12" x 6" FOV (part of MRS-L FOV)
  - 0".403 /pix

## Spec.: Low Resolution Spectrograph (LRS)

- Wide wavelength coverage

- High sensitivity

- LRS-S

- 5 -- 26  $\mu\text{m}$  covered by prism

- 2'.5 x 1".40 long slit

- $R \sim 50 - 100$

- LRS-L

- 20 -- 38  $\mu\text{m}$  prism or grating

- 2'.5 x 2".66 long slit

- $R \sim 50 - 100$

# LRS issues

- High sensitivity is required. -- Feasible
- Choice of disperser and wavelength coverage
  - Prism : Wide coverage, Low Resolution, fab. Risk
  - Grating: Limited coverage, very feasible
  - Instantaneous wave length coverage for LRS-S
    - current plan : as wide as possible 5 -- 26 $\mu$ m
    - option 5 --10 $\mu$ m ( outside MRS, competitive JWST) or 10 -- 20 $\mu$ m (inside MRS, outside JWST)
    - Slit width should be optimized at longest wavelength

## Extension to longer wavelength

- Detector technology up to 50 $\mu$ m is under development
- Risk:
  - Detector prop. is unknown
  - Easy to cancel the extension and back to Si:Sb based 38 $\mu$ m limit
- Cost:

# Expected performance

## ■ WFC-S/L

- Point source, 1hr,  $5\sigma$
- 4.9 / 6.7 / 7.7  $\mu\text{Jy}$  for 20 / 30 / 38  $\mu\text{m}$
- 0.13 / 1.2 / 3.5  $\mu\text{Jy}$  for 5 / 10 / 20  $\mu\text{m}$

## ■ HRS-S/L

- Point source, 1hr,  $5\sigma$  8mJy(L) 3mJy(S)
- Diffuse , 1hr,  $5\sigma$  10 MJy/sr(L) 30 MJy/sr(S)

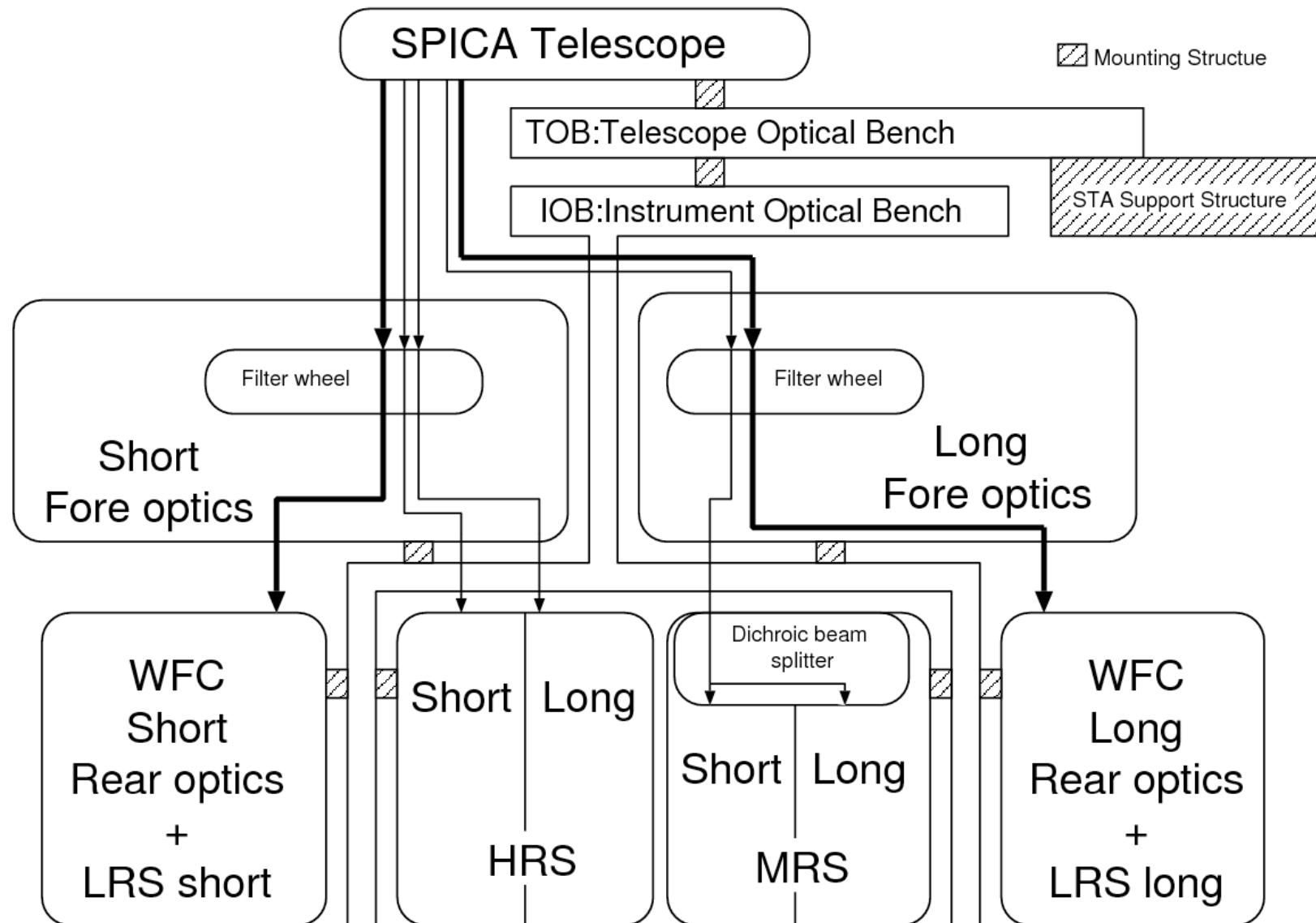
## ■ MRS-S/L

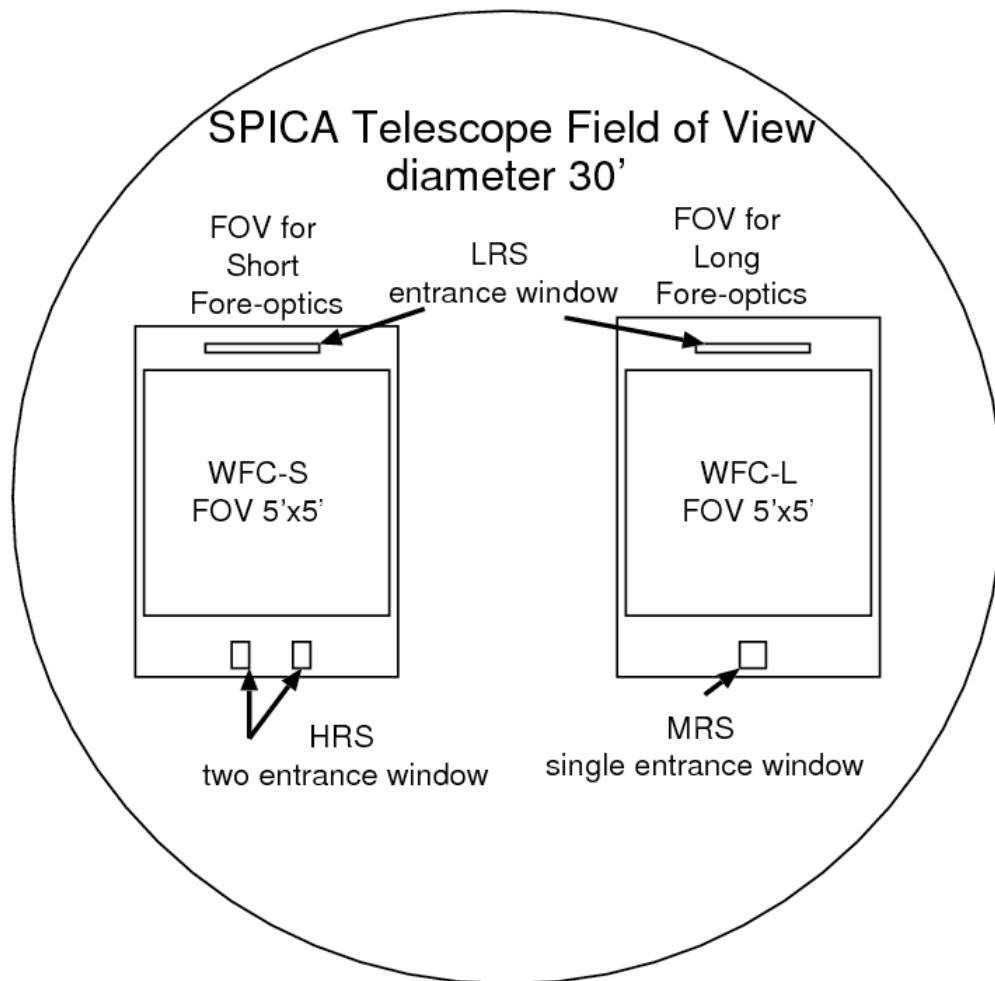
- Point source, 600sec,  $5\sigma$  ~600  $\mu\text{Jy}$ (L) ~100  $\mu\text{Jy}$ (S)

## ■ LRS-S/L

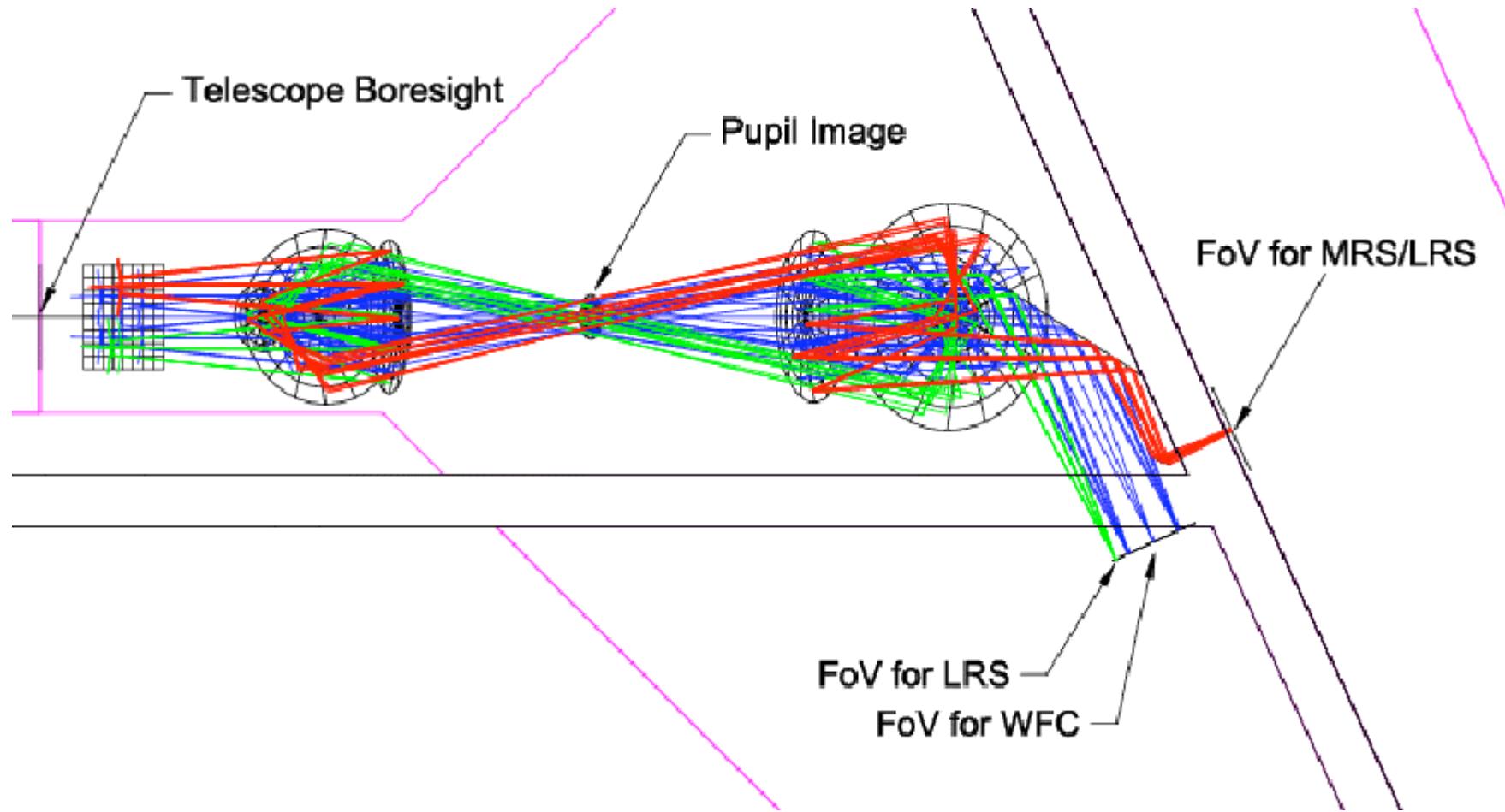
- Point source, 600sec,  $5\sigma$
- 64 / 84 / 108 / 136  $\mu\text{Jy}$  at 20 / 25 / 30 / 35  $\mu\text{m}$
- 7.3/14/26/39  $\mu\text{Jy}$  at 7.5 / 10 / 15 / 20  $\mu\text{m}$

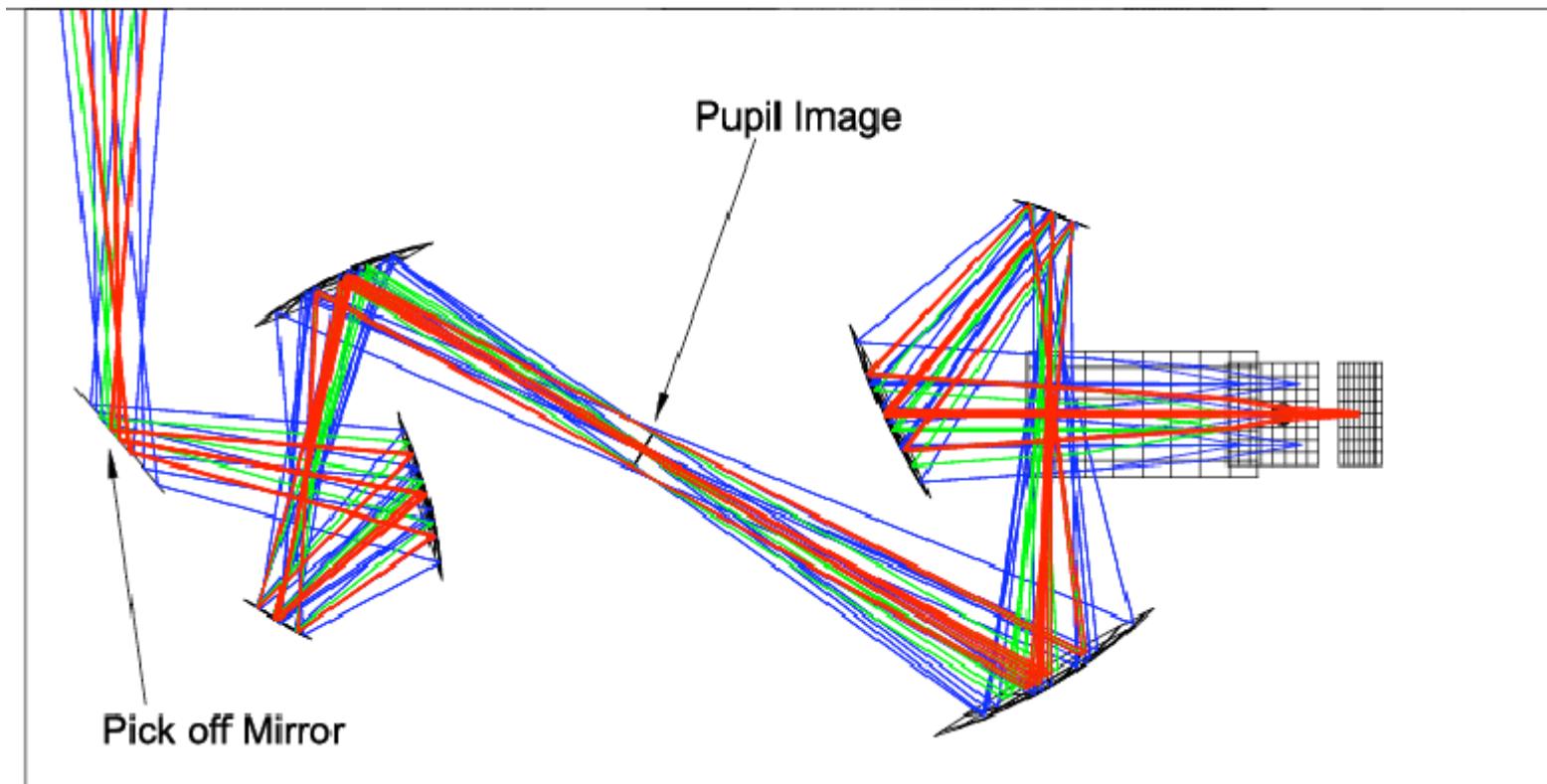
# Design: Optical architecture





# Design: Fore-optics

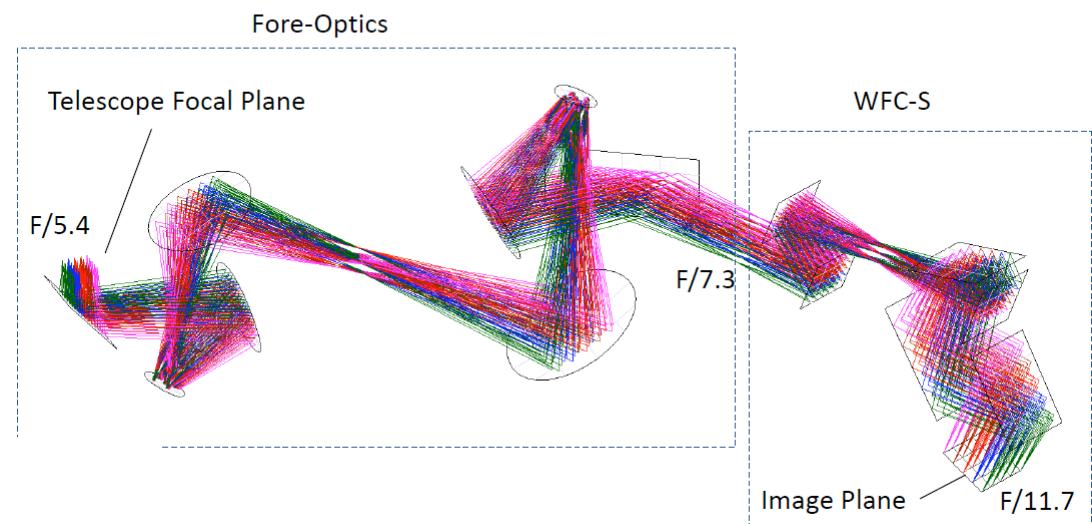




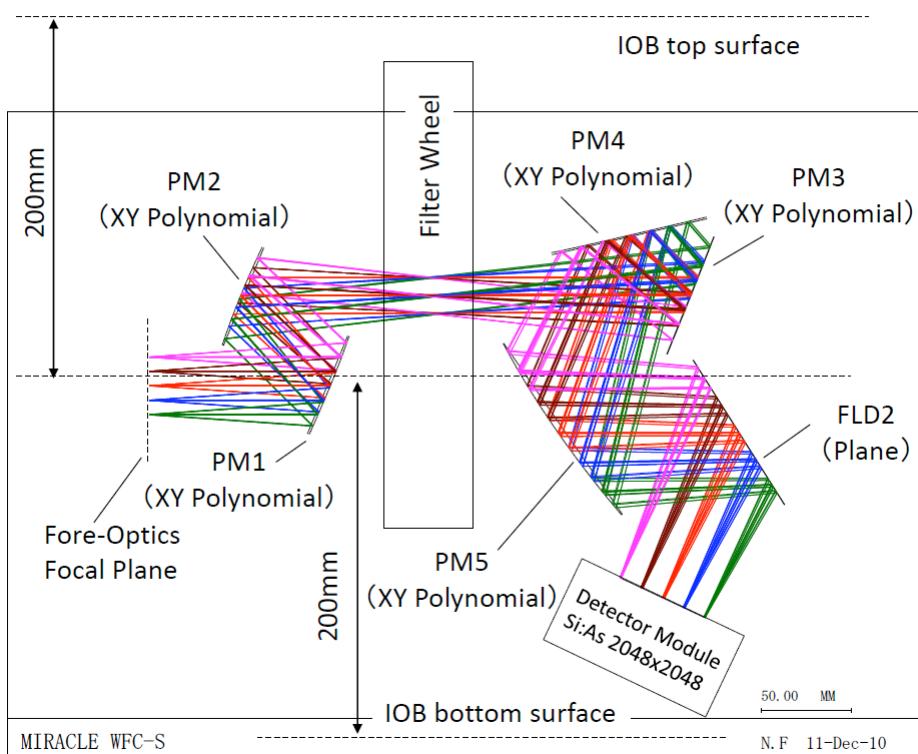
# WFC-S 3D View

## Design: WFC

### ■ WFC-S

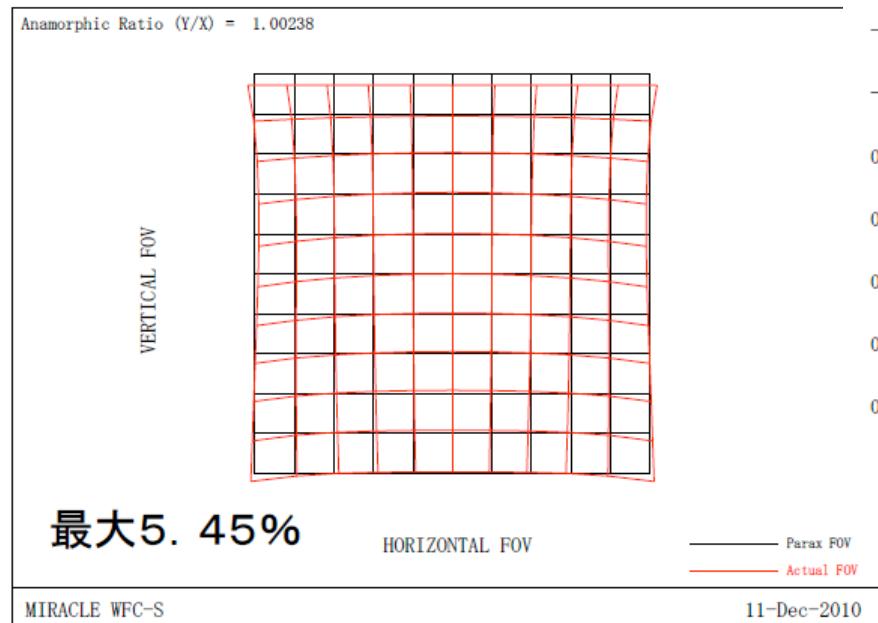


### YZ Cross-Section



# ■ WFC-S

## ディストーション



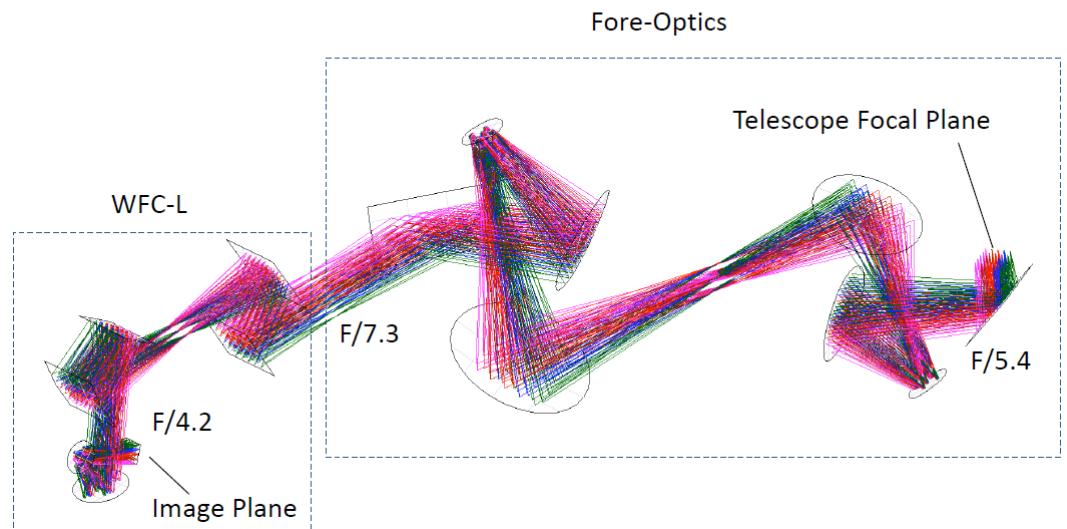
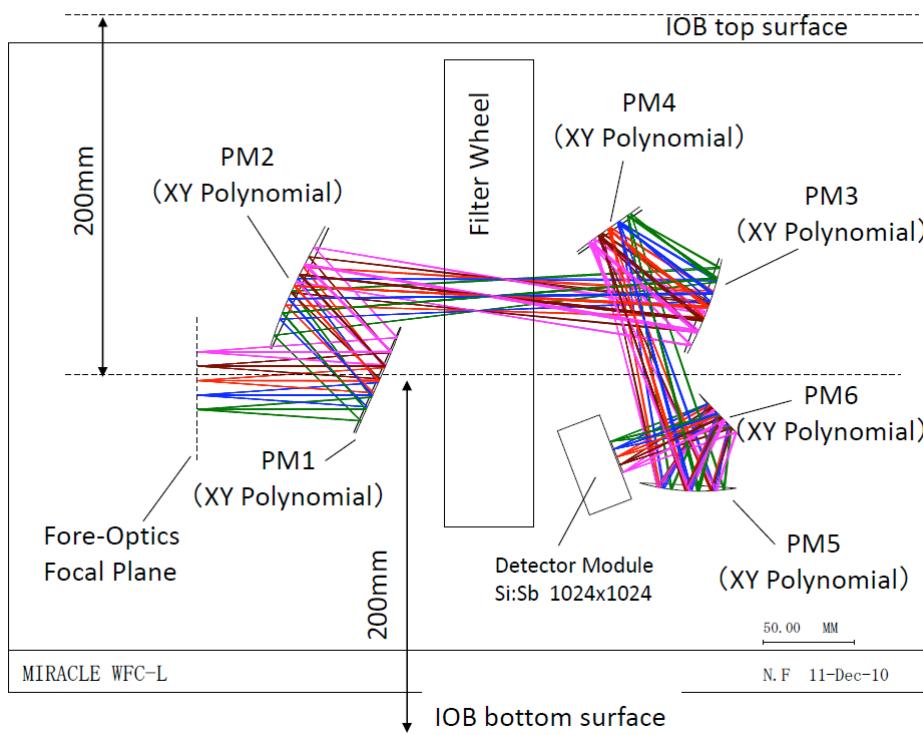
画角番号	ストレール比 @波長5um	
F15	0.969	RMS 11-Dec-2010 0.044705
	100% = 0.173109	
	RMS = 0.059006	
F14	0.905	100% = 0.186399
	RMS = 0.056750	
F13	0.897	100% = 0.151961
	RMS = 0.042918	
F12	0.969	100% = 0.153191
	RMS = 0.051423	
F11	0.884	100% = 0.143878
	RMS = 0.047629	
F10	0.937	100% = 0.139538
	RMS = 0.036731	
F9	0.958	100% = 0.093192
	RMS = 0.039009	
F8	0.948	100% = 0.101713
	RMS = 0.039766	
F7	0.978	100% = 0.146600
	RMS = 0.040494	
F6	0.952	100% = 0.081611
	RMS = 0.043095	
F5	0.955	100% = 0.120823
	RMS = 0.042700	
F4	0.926	100% = 0.077014
	RMS = 0.053277	
F3	0.914	100% = 0.106202
	RMS = 0.042539	
F2	0.959	100% = 0.121616
	RMS = 0.064531	
F1	0.889	100% = 0.112385



# WFC-L 3D View

## ■ WFC-L

### YZ Cross-Section

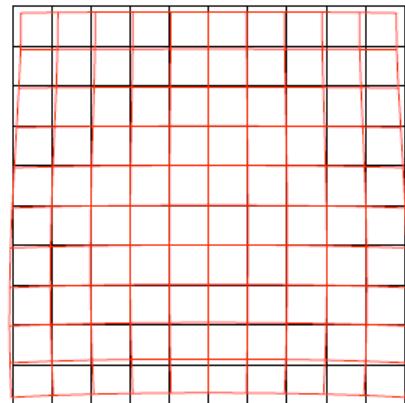


# ■ WFC-L

## ディストーション

Anamorphic Ratio (Y/X) = 1.00783

VERTICAL FOV

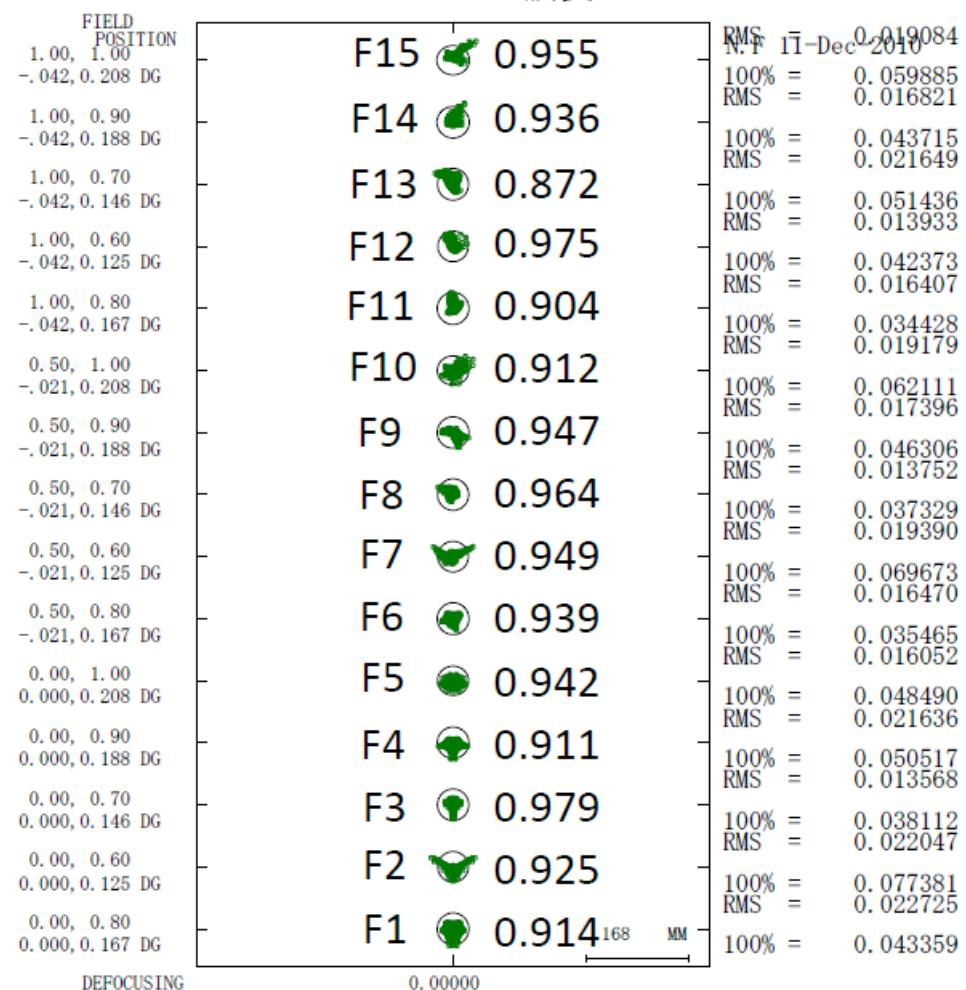


最大5.79%

HORIZONTAL FOV

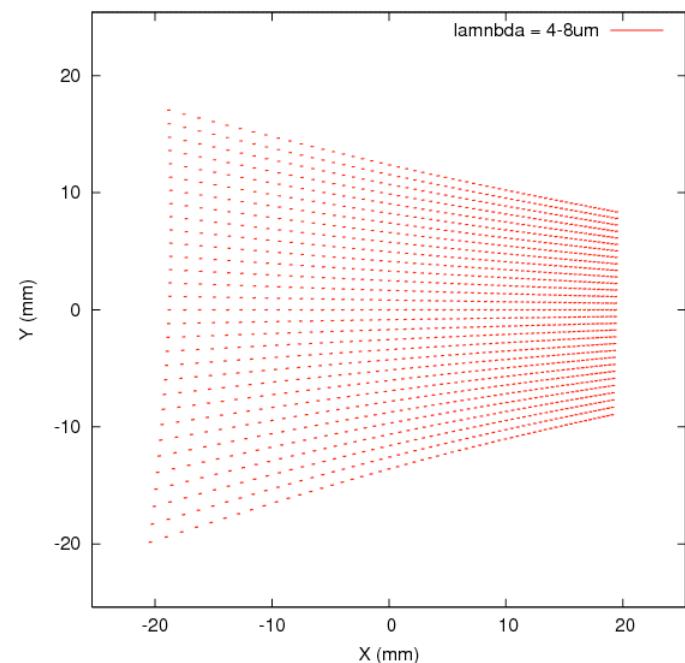
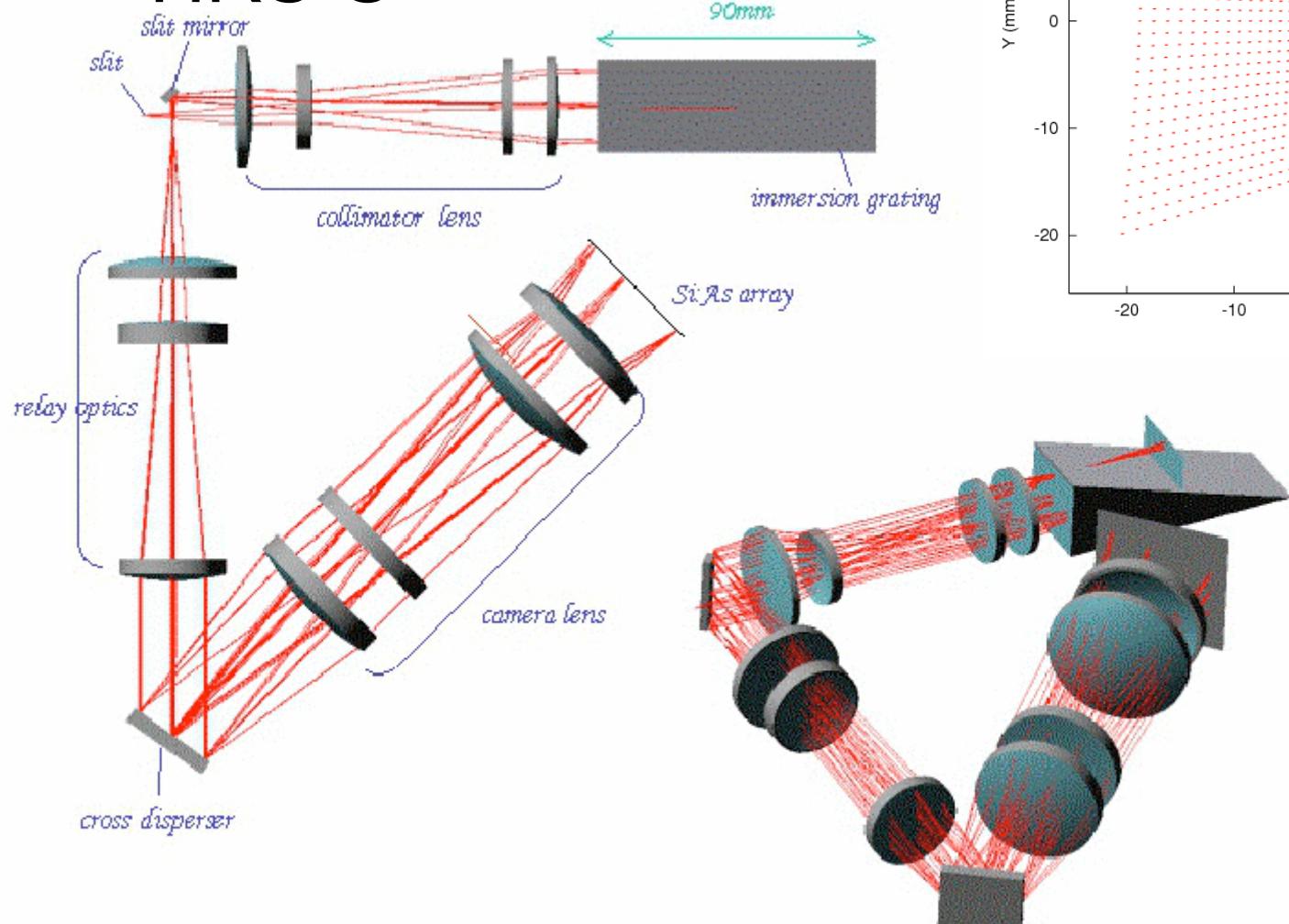
Parax FOV  
Actual FOV

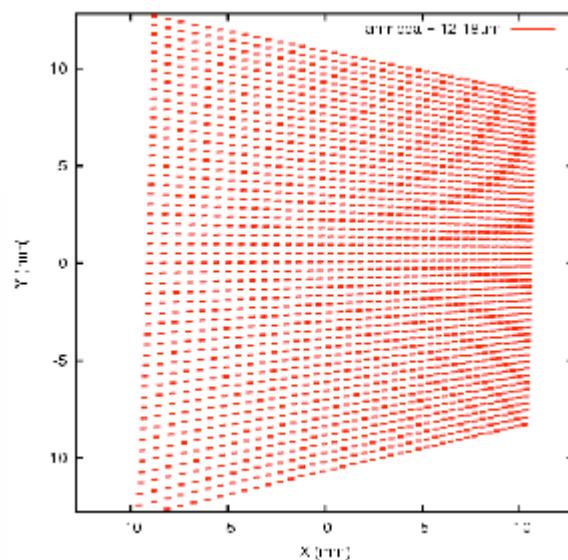
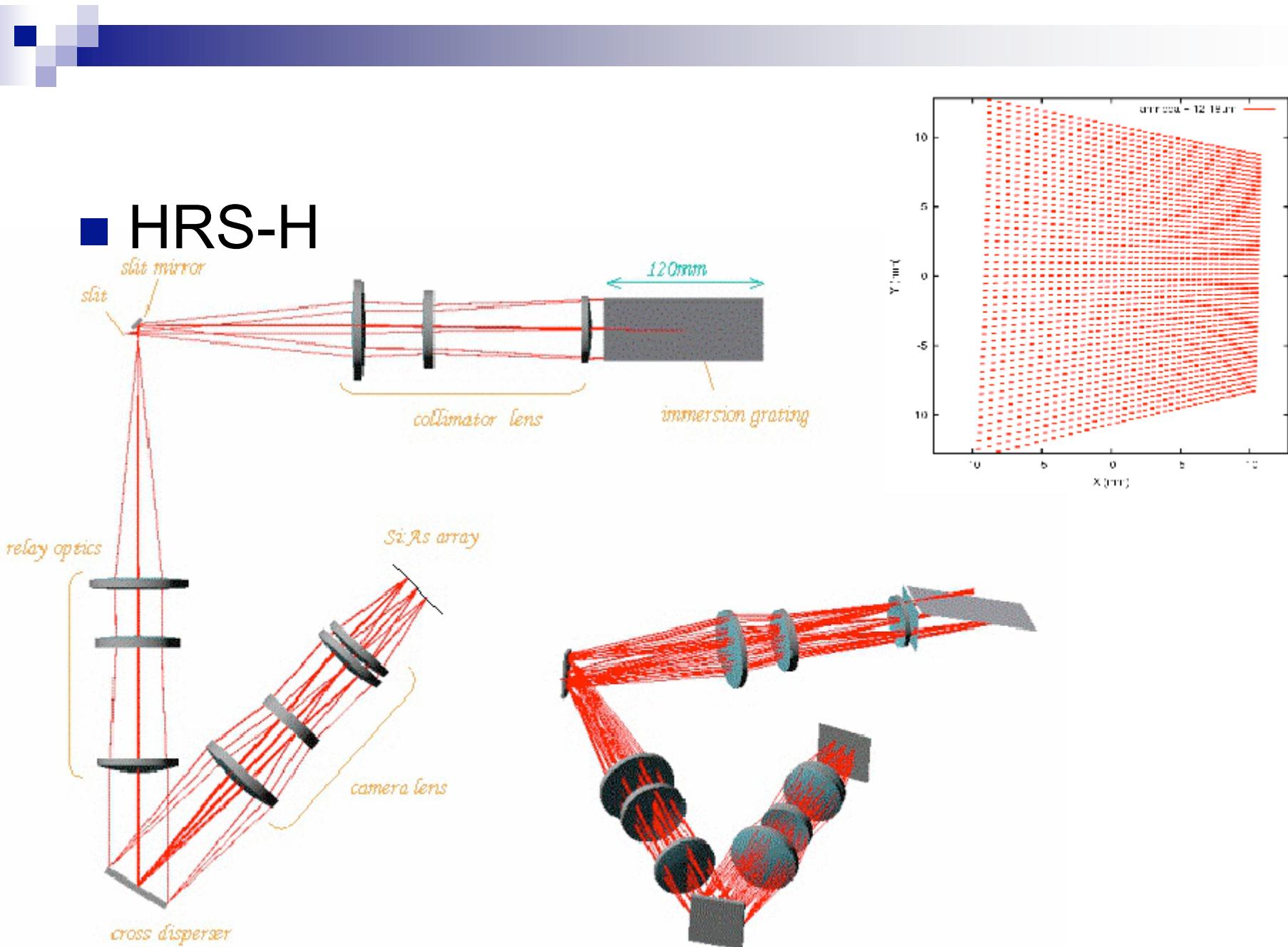
ストレール比  
画角番号 @波長5um



# Design: HRS

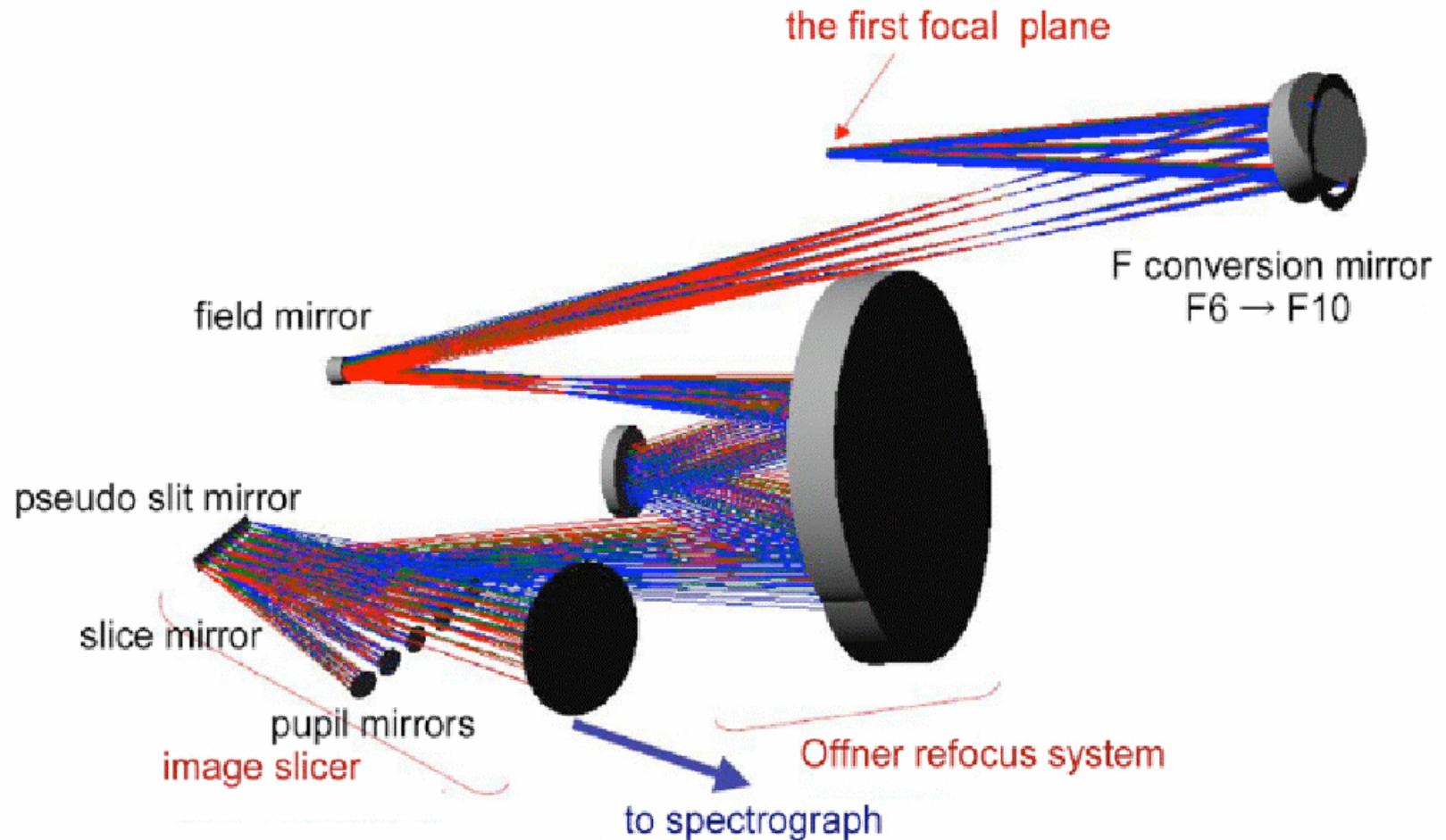
## HRS-S



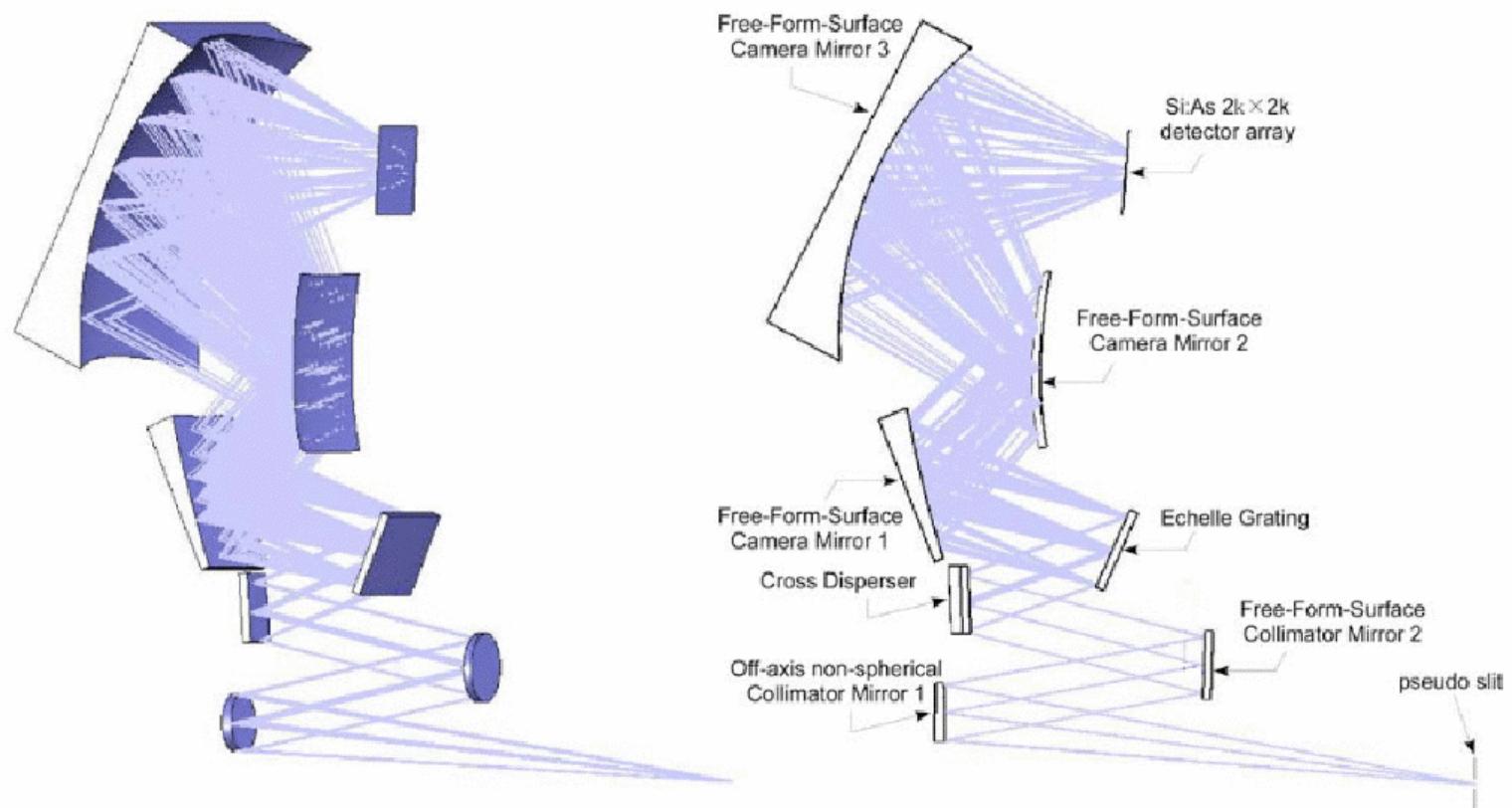
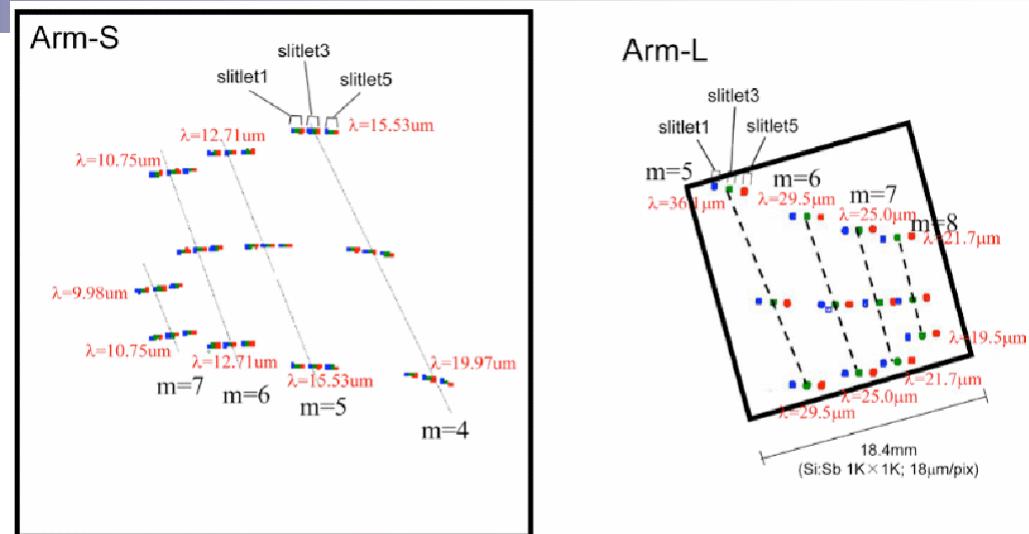


# Design: MRS

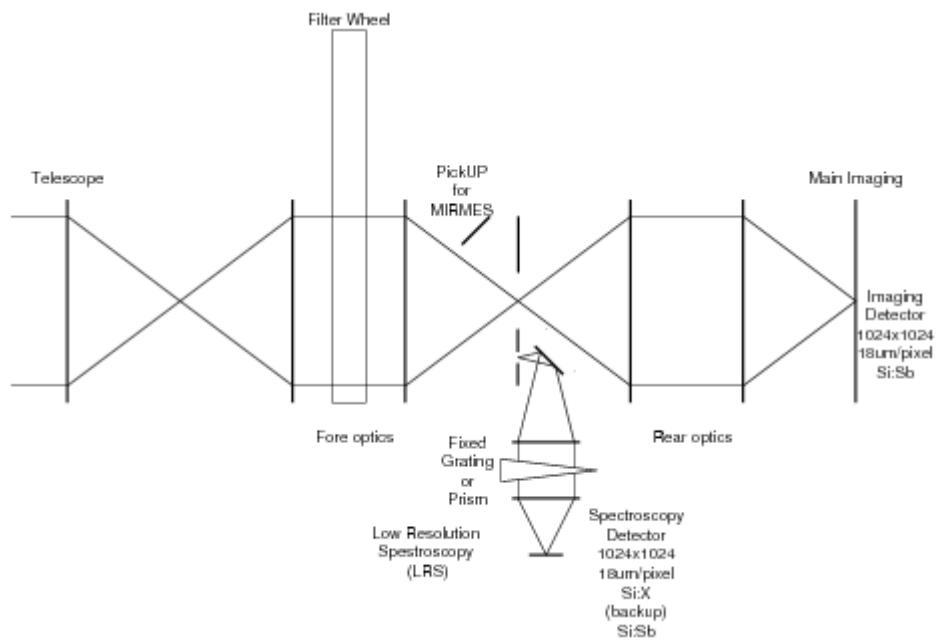
## ■ Slicer



# ■ spectrograph



# Design: LRS



# Design: Heat and Mass

## ■ Thermal design:

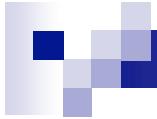
- VERY Limited resource
- Operation mode is limited, Clear the resource limit

## ■ Mass budget:

- Rough estimate, Within the budget

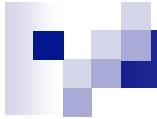
Table 20: Operation mode and heat load

Mode	20K (mW)	4.5K (mW)	1.7K (mW)	note
Standby Allocated	TBD	0.4	0.3	
Observing Allocated	TBD	4.0	2.1	
Standby	2.24	0.18	<0.01	
WFC-S/L + LRS-S/L	13.4	3.6	2.0	Frame rate 1Hz(WFC-L,LRS-L), 1/4Hz(WFC-S,LRS-S)
WFC-S/L + MRS-S/L	13.4	3.6	2.0	Frame rate 1Hz(WFC-L,MRS-L), 1/4Hz(WFC-S,MRS-S)
WFC-S/L + HRS-S/L	10.64	3.6	1.0	Frame rate 1Hz(WFC-L), 1/4Hz(WFC-S), 1/8Hz(HRS-S/L)



## Devel.: Detector

- Si:Sb 1024 x 1024
  - On going development. Budget limited.
- Si:As 2048 x 2048
  - JWST Heritage, Start to develop.
  - Budget limited
- Si:As High dope 1024x1024
  - Development by another US project. (We can get information. )



## Devel.: Optical elements

### ■ Immersion Grating

- HRS-S : ZnSe final development phase
- HRS-L : CdTe Now, look hopeful

### ■ Filter, Prism, Mirror, Lens, and etc

- On going development.
- Man power / technology / budget limit