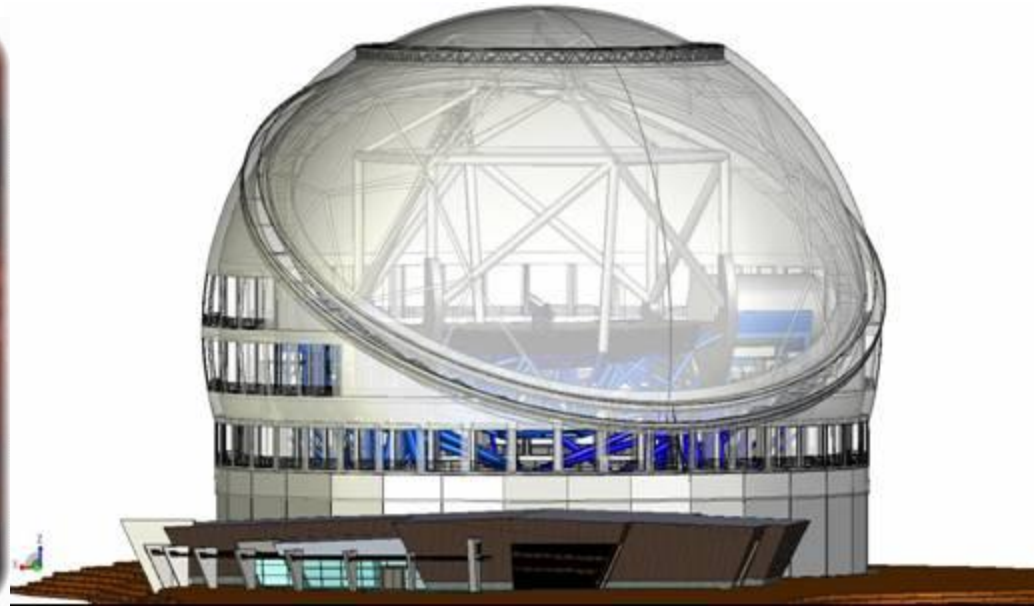
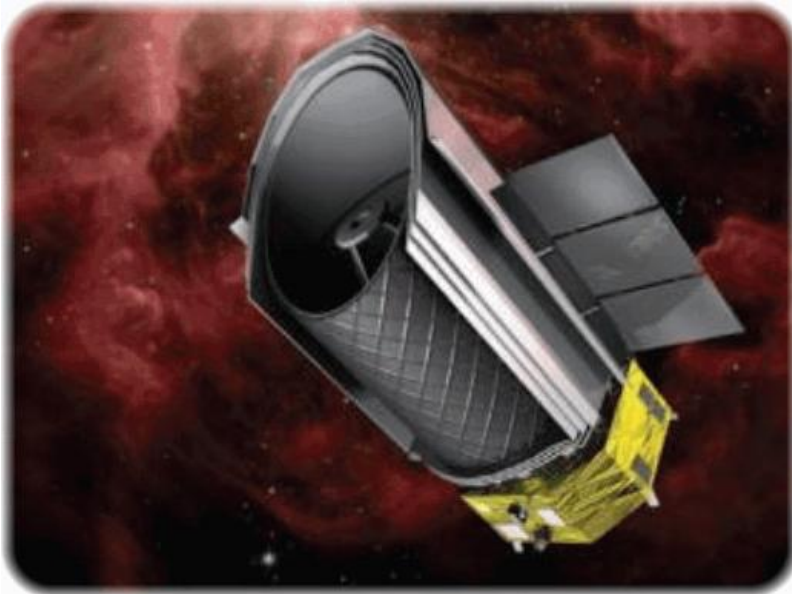


Synergy of SPICA and TMT



- Congratulations for the SPICA community having this many participants with very intensive and candid science discussion!

- **But, schedule conflict with Subaru autumn school ...**

Masanori Iye (TMT-J project)

Thirty Meter Telescope Project

- 492 hexagonal mirrors of 1.5m, 30m effective array
- Costs ~1300\$ (Secured ~30%), FL expected in 2019
- Three FL instruments: IRIS, WFOS, IRMS
- TMT board meets quarterly in Pasadena
- Partners: Caltech+ U.California+ ACURA+ NAOJ+NAOC+ India+ NSF
- NSF decision expected at the end of 2011.
- Science Council of Japan recommendations



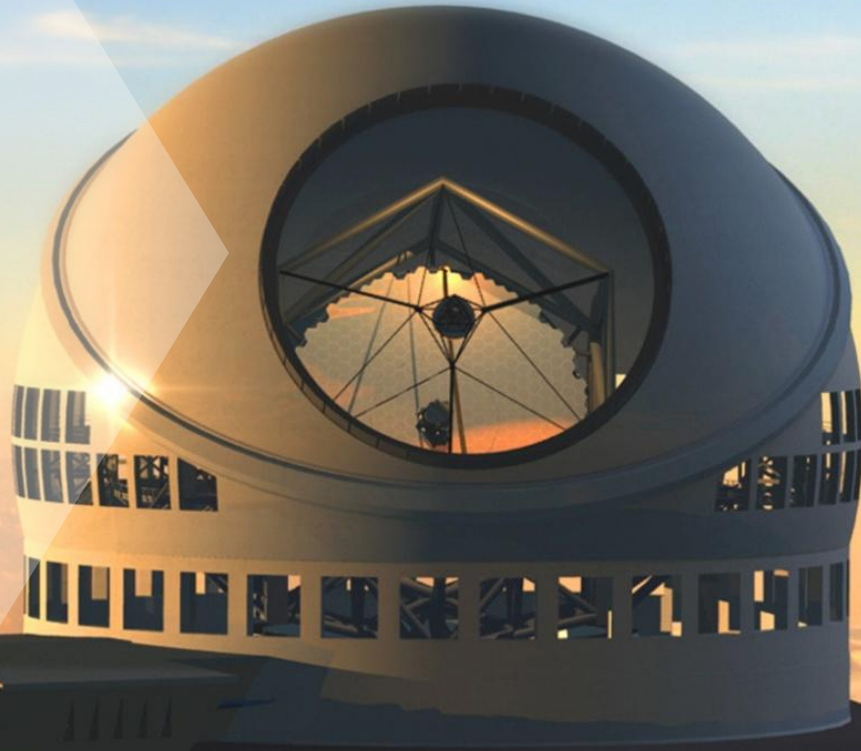
“TMT-J Status Report”, M.Iye
(NAOJ), EAMA, Shanghei,

101013

TMTからSPICAへの期待

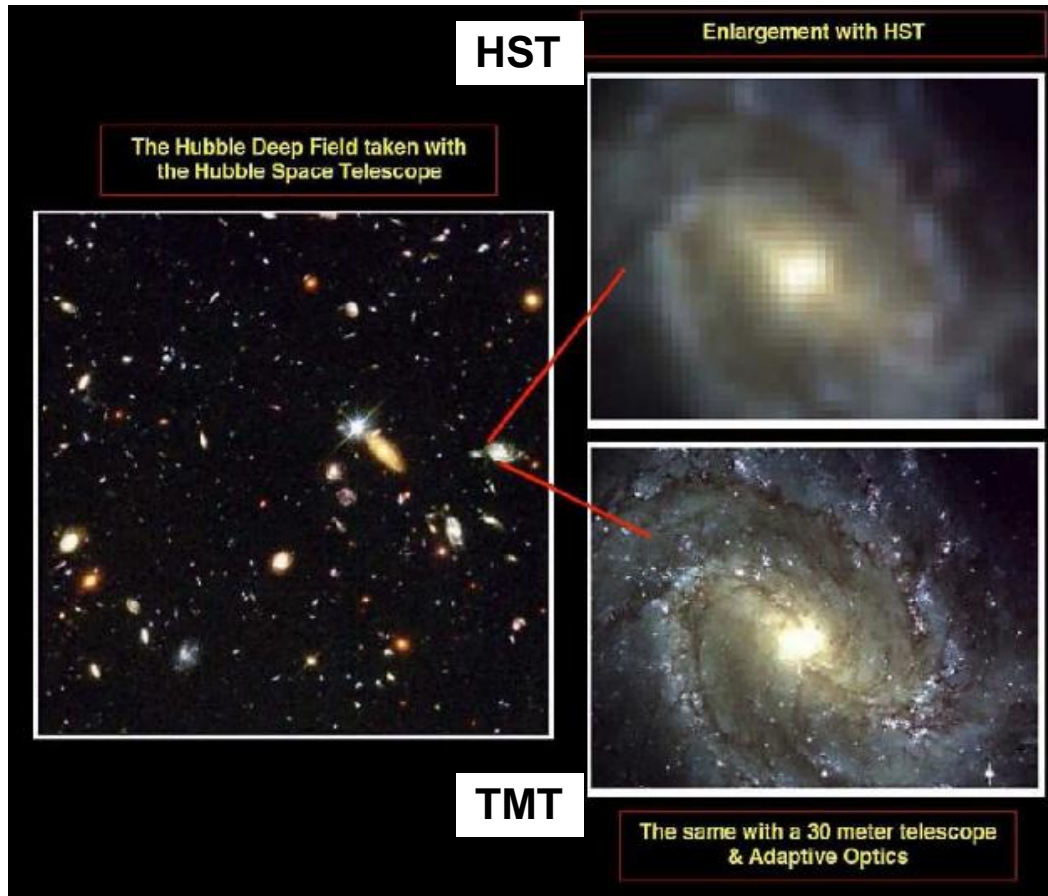
The Thirty Meter Telescope Project

- ◆ Overview
 - ◆ Telescope
 - ◆ Instruments
 - ◆ Site
 - ◆ Japanese Contributions
- ◆ TMT Science Cases
- ◆ Synergy w/SPICA



Nobunari Kashikawa
(NAOJ/TMT project) June 2009

Adaptive Optics(AO)



- Angular resolution $0.6 \rightarrow 0.015$ arcsec ($2.2 \mu\text{m}$)
- sensitivity upgrades by 1 order
- Several thousand elements
- Much higher resolution than HST
- 5 times higher resolution than JWST
- Almost all the TMT NIR observation will use AO in TMT.

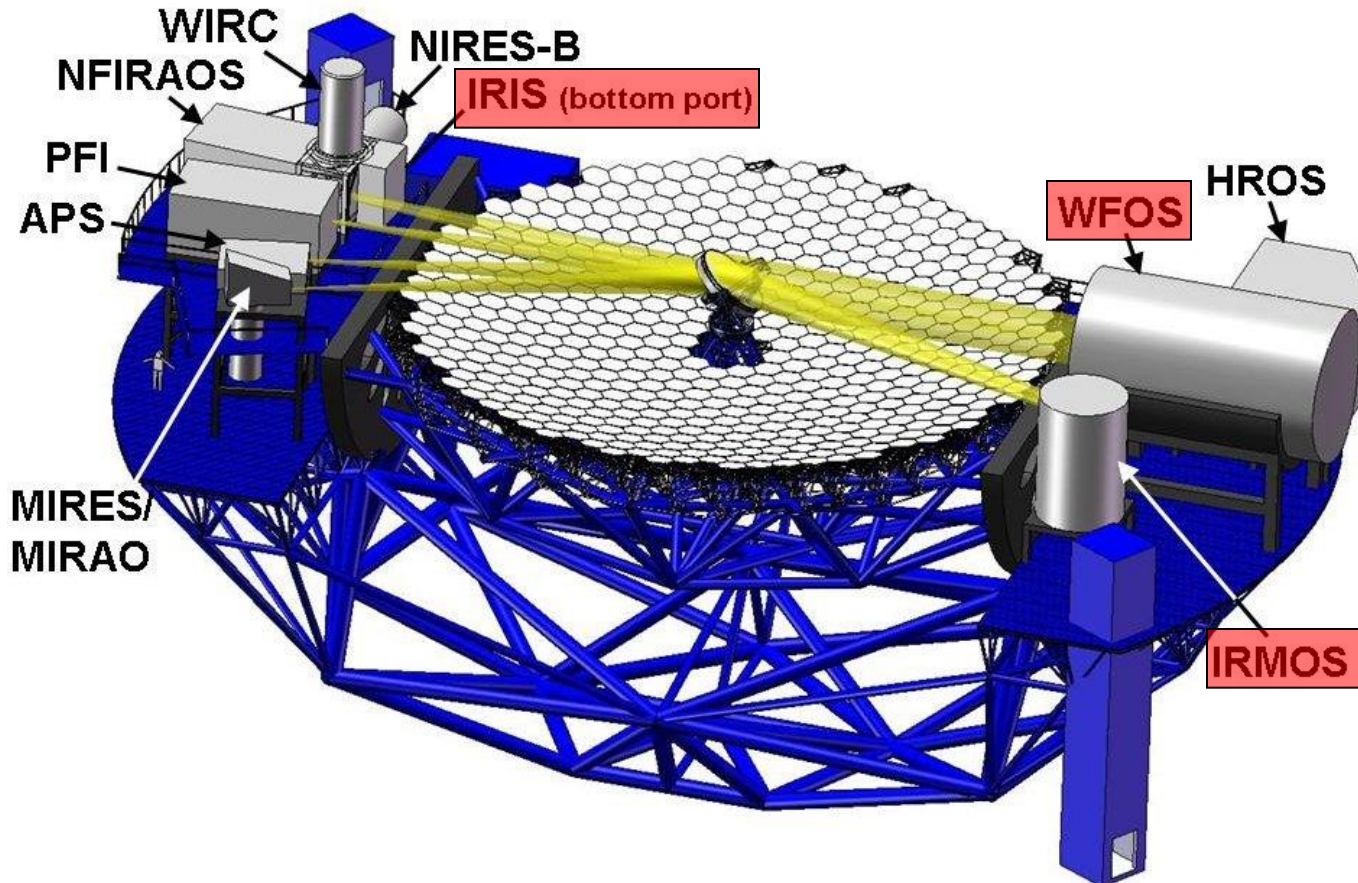
Resolution & Sensitivity

	Resolution @ $\lambda=10\mu\text{m}$	Sensitivity 5σ 1hr @ $\lambda=10\mu\text{m}$	Spectr.Resolution R
30mTMT + MIR	0.09"	150 μJy	数百、数千、100,000
8mSubaru + COMICS	0.32"	2,000 μJy	250 2,500 10,000
3.5mSPICA + MIR	0.75"	1 μJy	1,500 3,000 (30,000)
6.5mJWST + MIRI	0.4"	0.1 μJy	3,000

- 宇宙望遠鏡の1桁上をいく空間分解能
- 2桁上をいく波長分解能

Nasmyth Configuration: First Decade Instrument Suite

- ◆ Platform 7 m below elevation axis
- ◆ Articulated M3 – facilitates quick instrument change
- ◆ Addressable regions: -28° to 6° and 174° to 208° for small FOV



4. TMT-J Four Groups for Instruments

High Priority & Feasibility

- ◆ **MICHI (Mid-IR Camera, High-disperser, and IFU)** (Y.Okamoto+): Modified MIRES ← **COMICS, TAO**
- ◆ **NIR High Dispersion Spectrograph** (N.Kobayashi+): NIRES-b and NIRES-r ← **IRCS-HRU, WINERED & IRDI (Immersion Gr., Astro-Comb, fiber)**

Medium Priority & Feasibility

- ◆ **NIR Multi IFU spectrograph with MOAO** (M.Akiyama+): IRMOS ← **RAVEN, SIRMOS**
- ◆ **Optical High Dispersion Spectrograph** (W.Aoki+): HROS ← **HDS, HIDES (I2 cell, image slicer, fiber)**



• Big Objectives for next Decade

THIRTY METER TELESCOPE

★ First stars(Cosmic dawn)

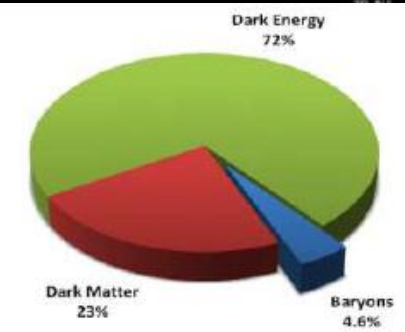
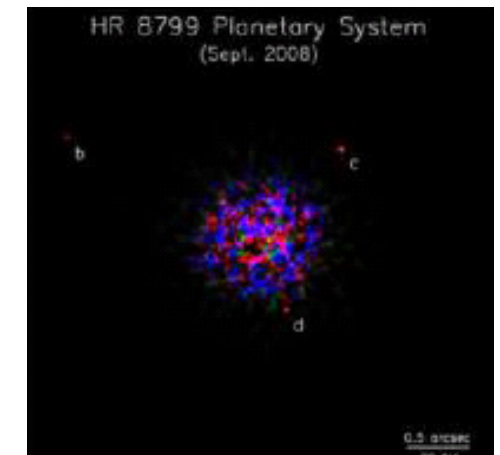
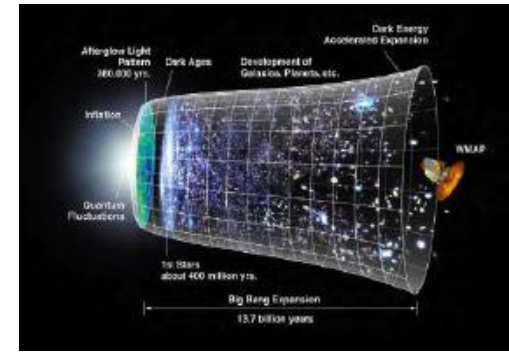
- ★ Look for first stars
- ★ Cosmic re-ionization

★ Second Earth

- ★ Habitable planets

★ Dark Energy

- ★ Cosmic expansion history
- ★ Modification for General relativity



Dark Energy

Direct measurement of cosmic expansion history (Feasibility assessment TBD)

•13Gyr

7Gyr

1Gyr

•Redshift



↓•10 yrs later

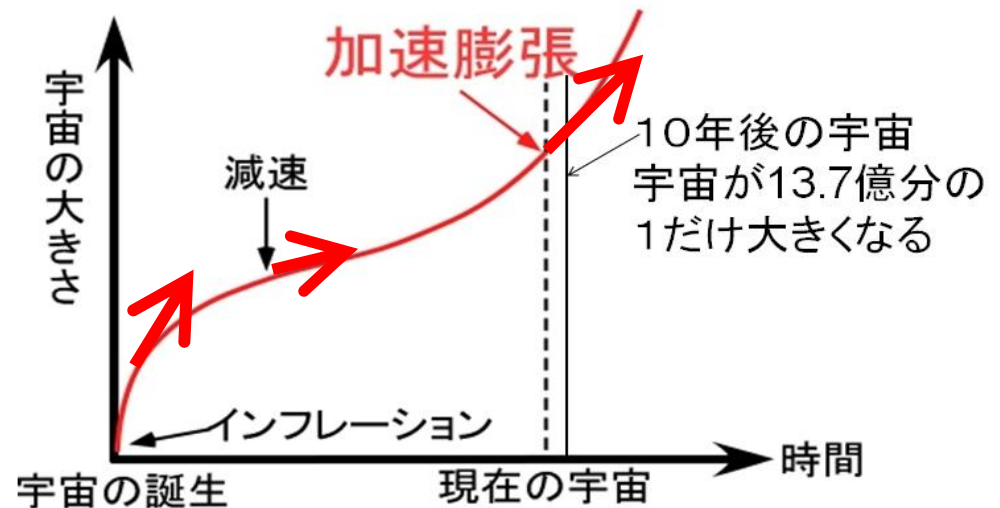


•Redshift

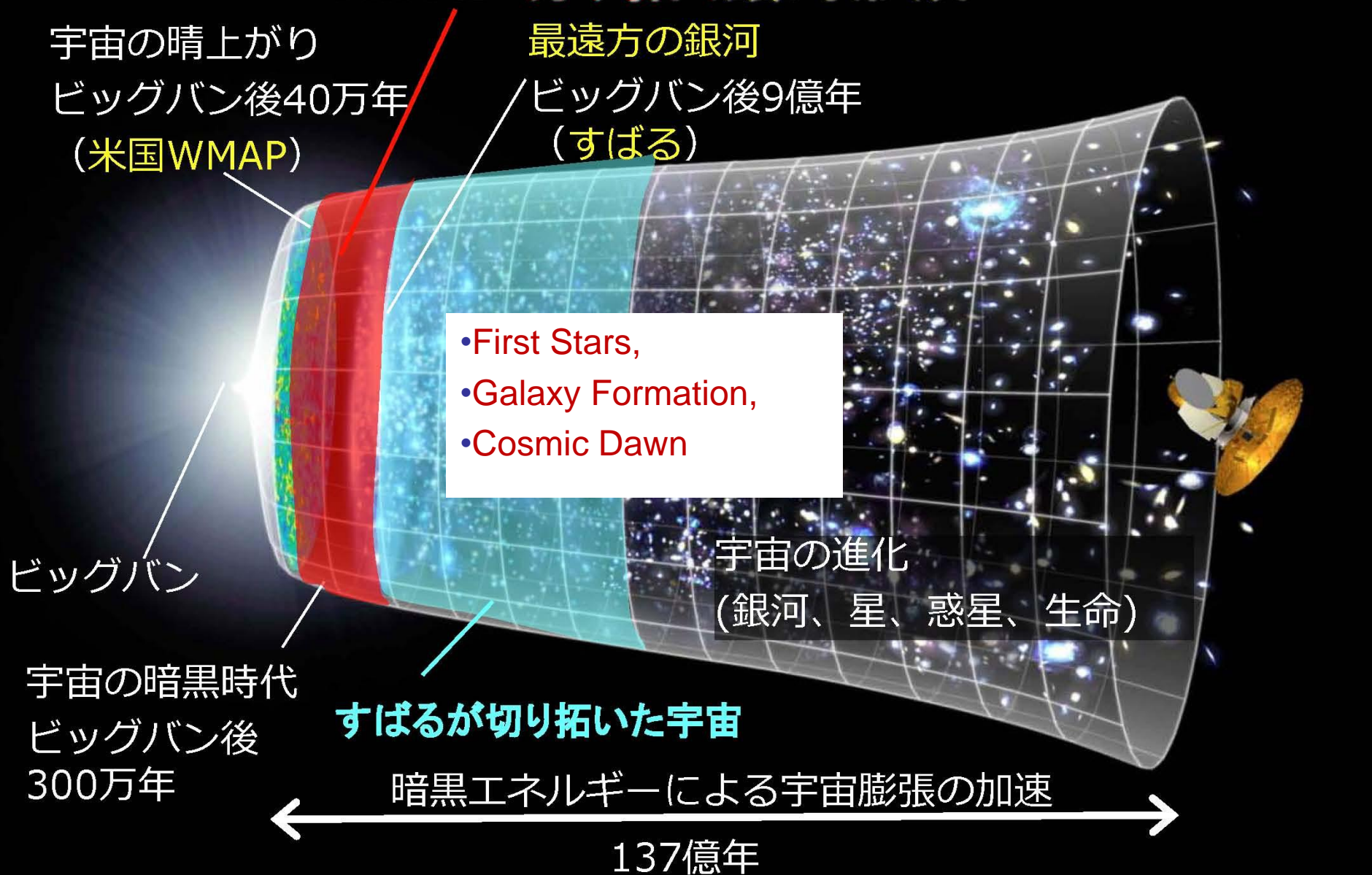
•Needs to measure variation of
10cm/sec over 10 years

•Doppler survey for exoplanets at
50cm/sec

•on 4m class telescopes

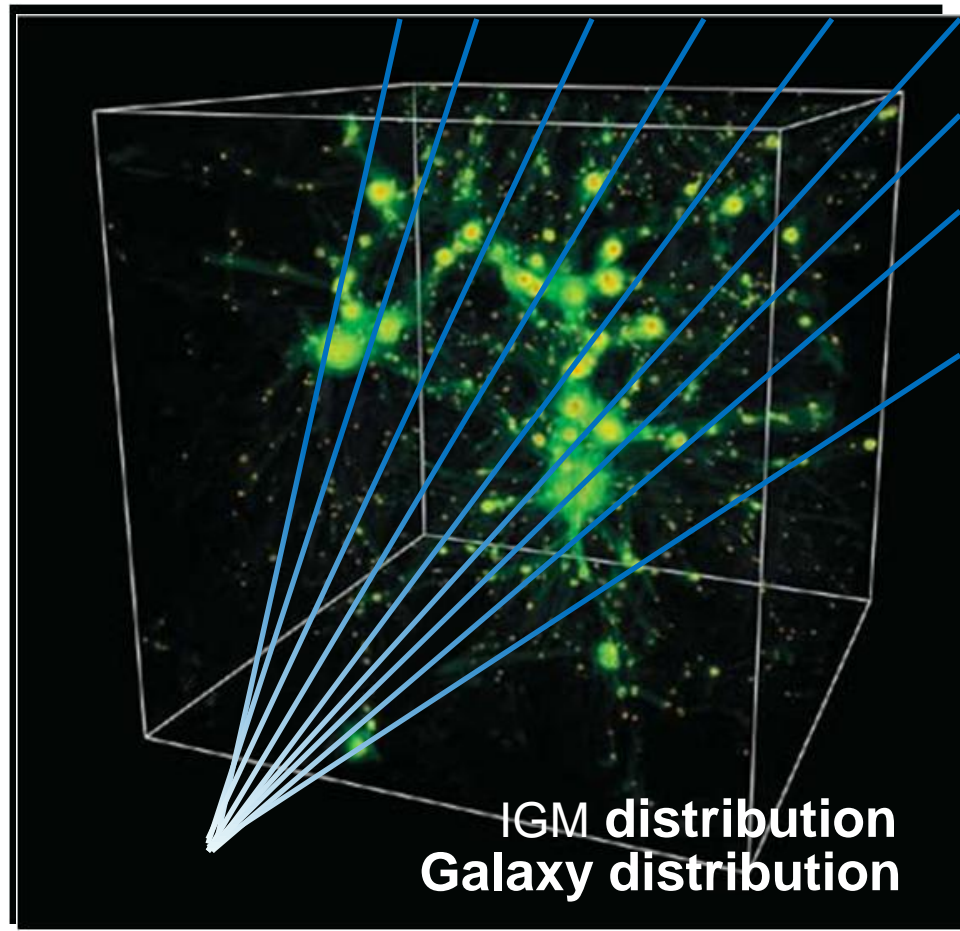
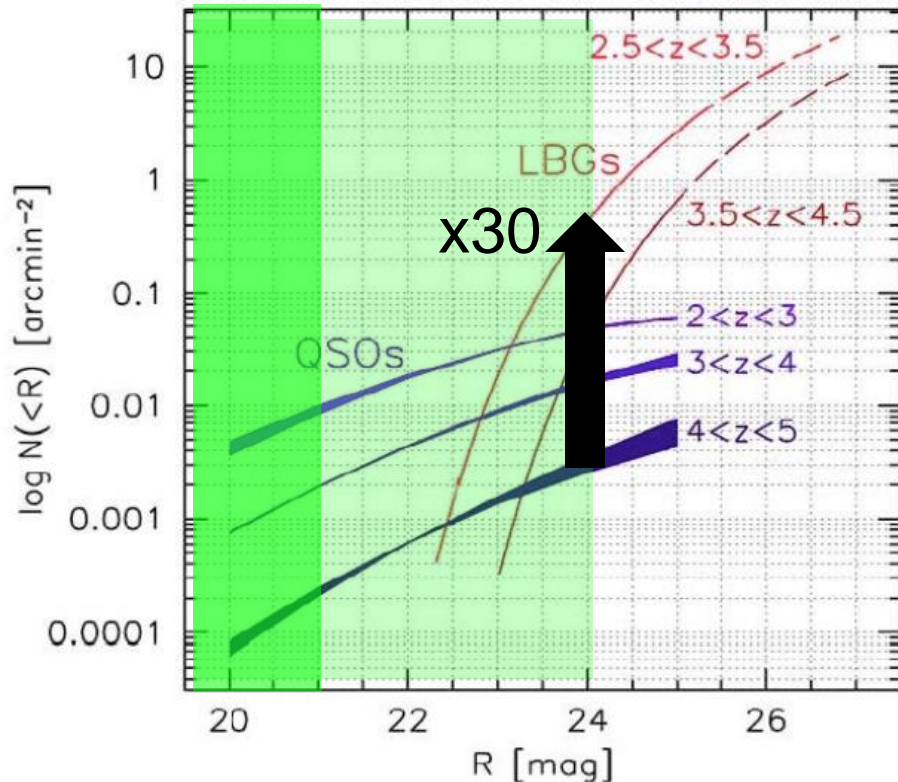


TMTが切り拓く銀河形成



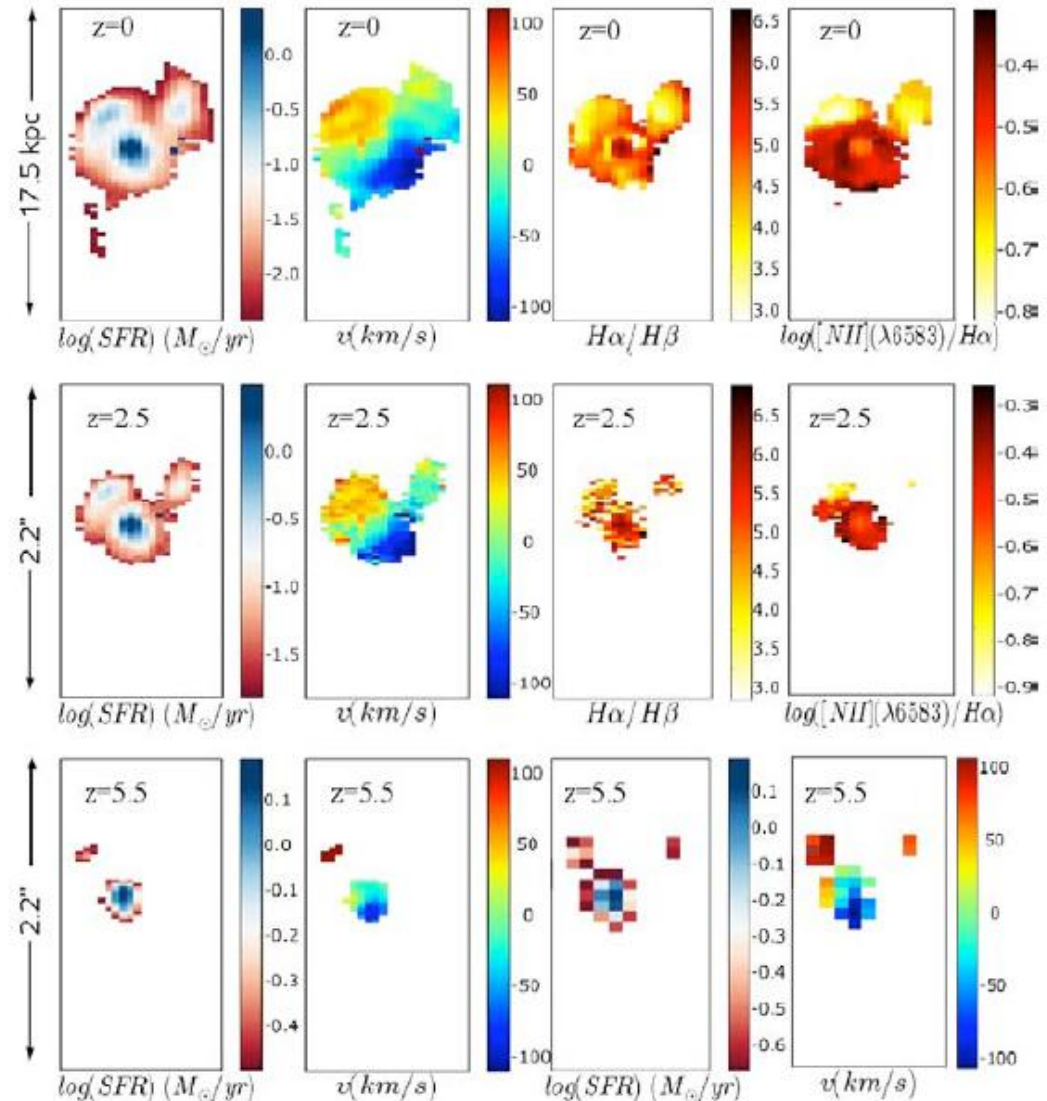
IGM tomography

- ◆ TMT R=10,000 mode: lim.mag.=24mag
- ◆ Not QSO but Galaxies are dominant in number density (2/arcmin²)
- ◆ Space correlation <300kpc scale
- ◆ 3Dmap of HI • metal • star • DM



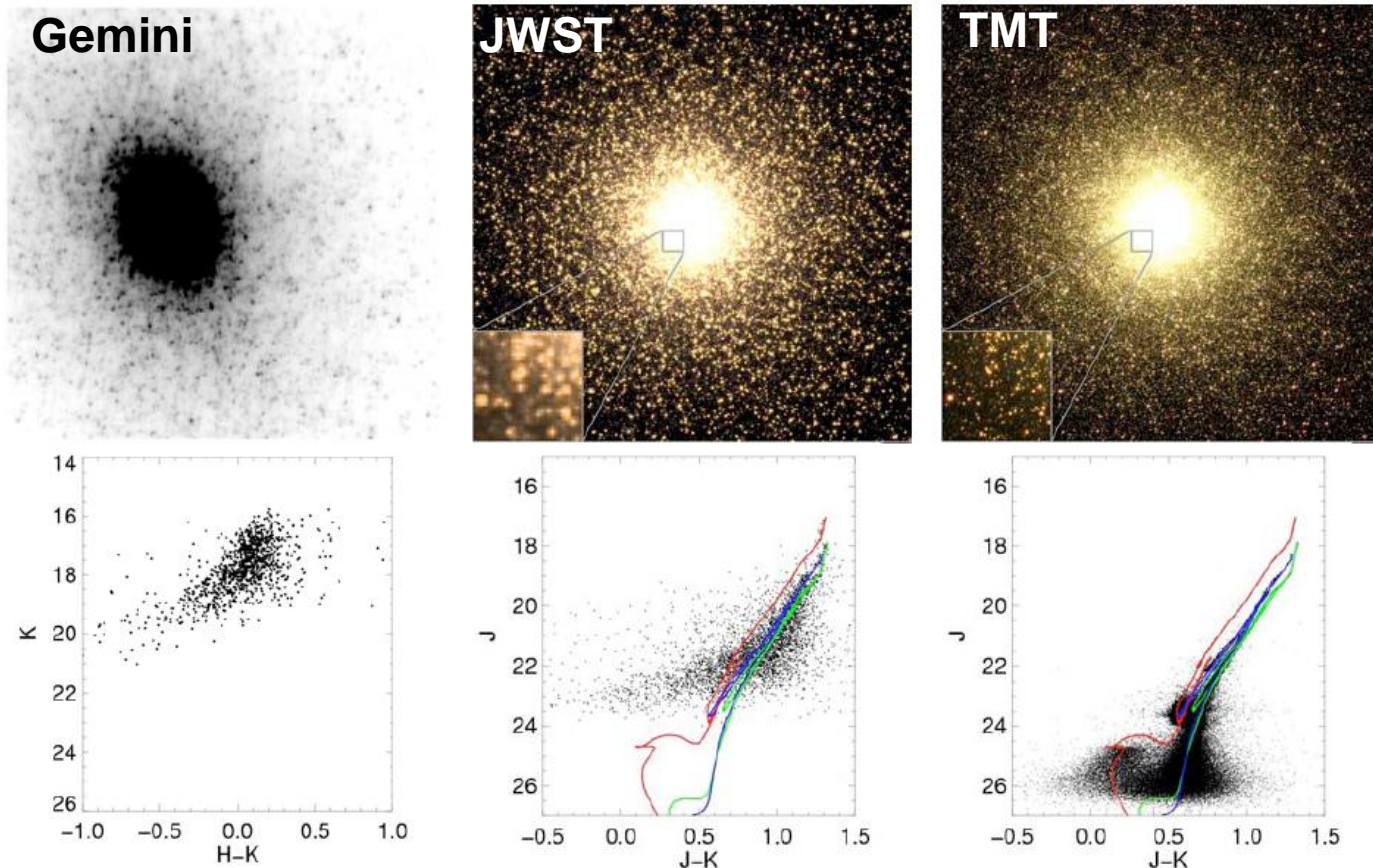
Mapping of kinematic/chemical evolutions

- 2D mapping of SFR, velocity, extinction, and metallicity.
- TMT will gain 10-100 in sensitivity and >3-5 in angular resolution over current facilities.
- Understand the Internal dynamics and complex baryonic processes within a DM halo.
- SPICA: dust, obscured AGN @ $z < 3$
- ALMA: molecular gas



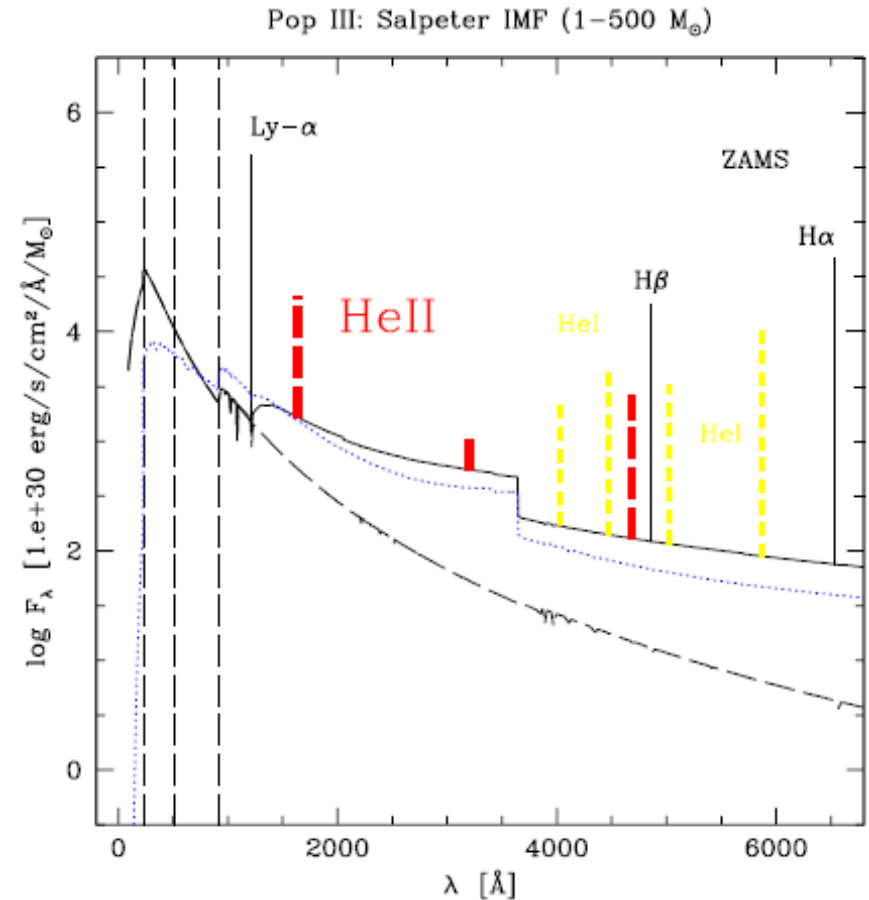
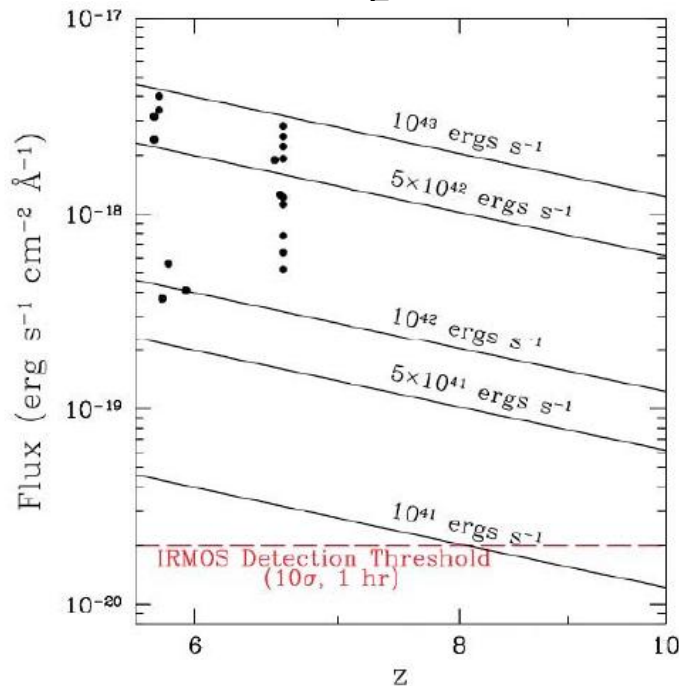
SFH of nearby galaxies

- ◆ Deep CMD for nearby galaxies can be drawn even in crowded regions. In the case of M32, TMT will reach to the MS turnoff.
- ◆ TMT R~4000 spec. for RGB stars will constrain SFH.
- ◆ SPICA: dust distribution, feedback from activity, merging



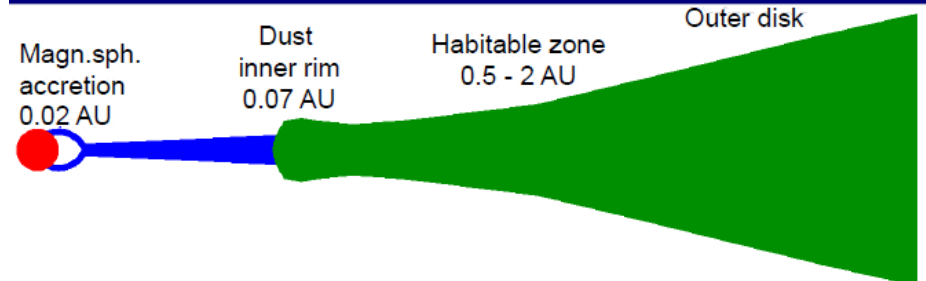
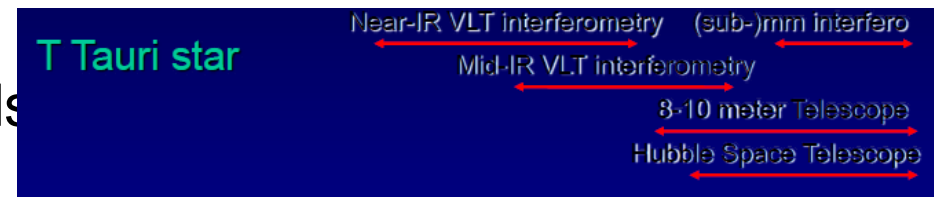
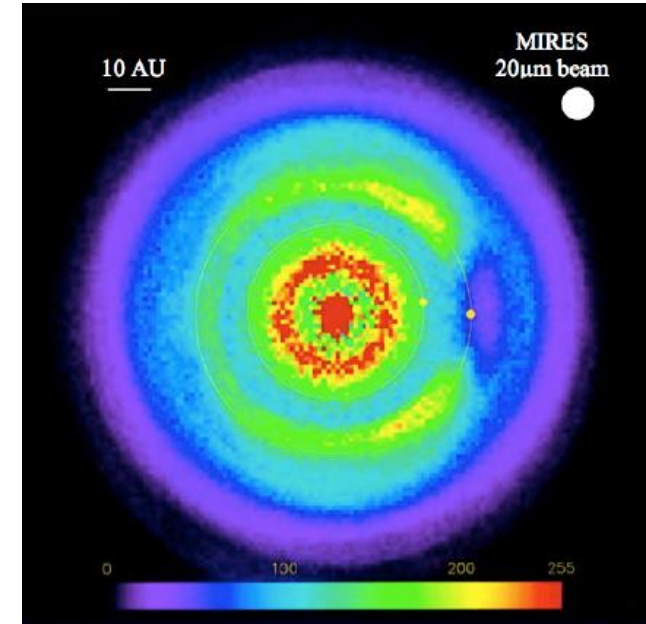
PopIII detection

- ◆ HeII 1640A is a characteristic signature of popIII.
- ◆ Tiny (<30mas) & faint sources
- ◆ TMT can detect HeII at $z < 14$.
- ◆ JWST: detection of sources
- ◆ SPICA: $H\alpha$, H_2 detection

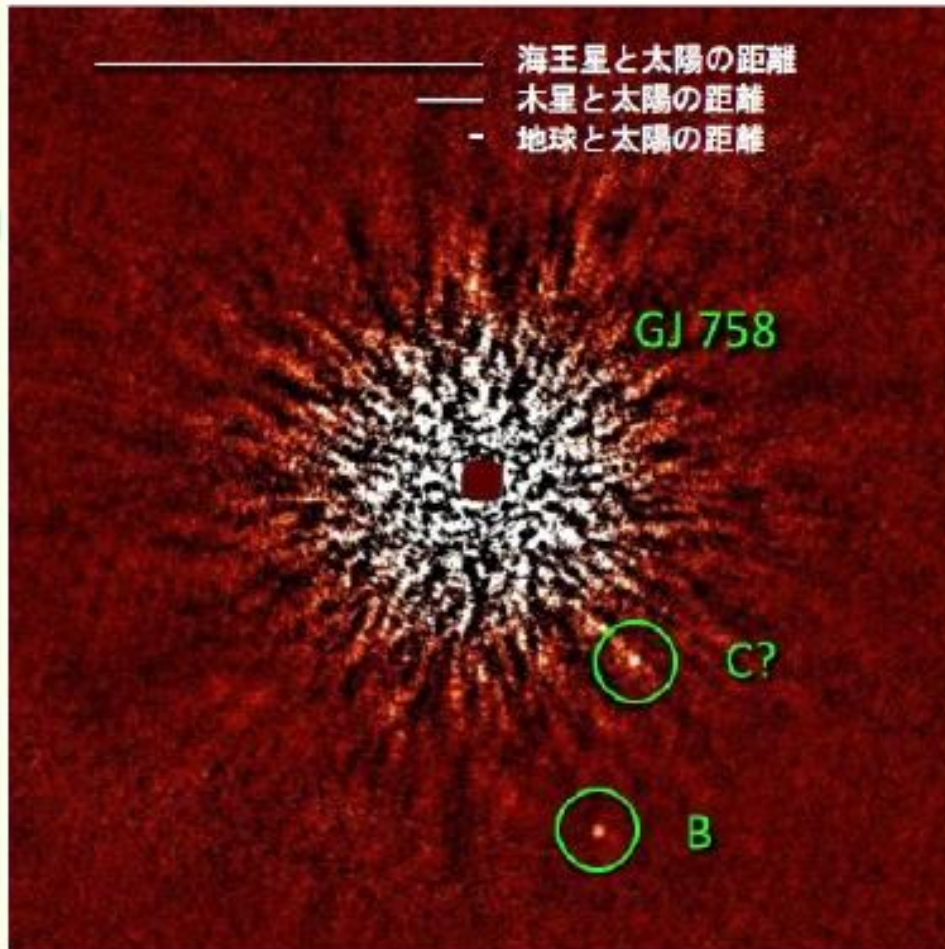


Protoplanetary disk

- ◆ TMT MIR R~100000 spec. for protostellar cores to reveal structure and kinematics of infalling envelopes /jets/winds.
- ◆ H₂O, CO as diagnostics to map T/ρ/v at <1AU
- ◆ MIRES will be able to image protoplanetary disks at <1au
- ◆ SPICA: H₂ flux, H₂O ice
- ◆ ALMA: outer molecular clouds



- Earth like Exoplanets in habitable zone
- Spectroscopy of atmosphere by TMT



惑星候補の性質

● 明るさ

性質	GJ758B	GJ758C
赤外線等級	19.3	18.5
距離（角度）	1.9"	1.2"
距離（AU）	29AU～海王星	18AU～天王星

● 質量と温度（明るさから求めた値）

年齢	GJ758B	GJ758C
7億年 （ベスト）	10木星質量 270℃	12木星質量 360℃
87億年 （最大値）	40木星質量 360℃	41木星質量 460℃

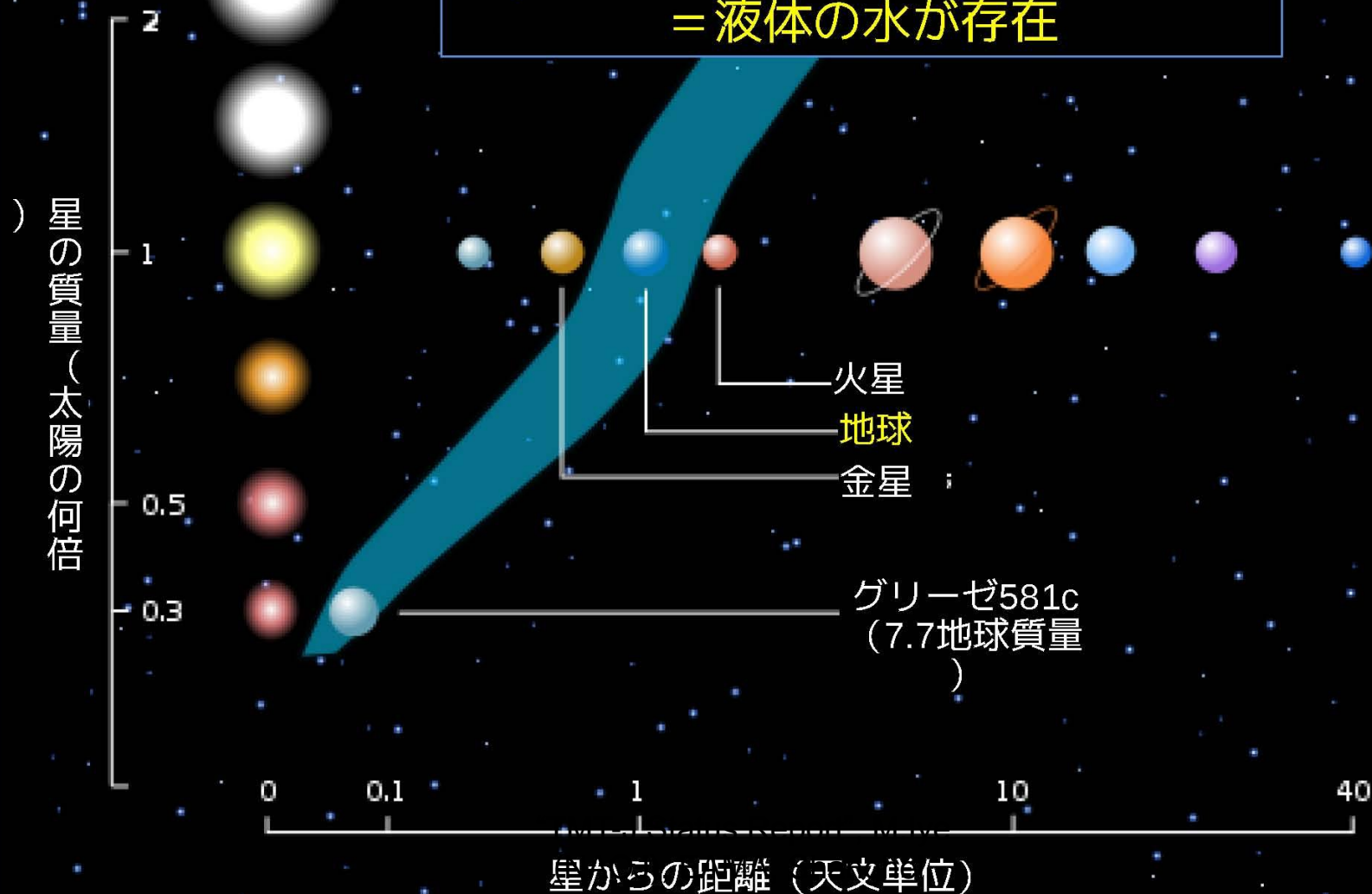
標準モデル、あるいは、
 重力不安定性でも説明難
 ⇒ 惑星散乱？

Exoplanets in habitable zone

Spectroscopy to search for biomarkers in the atmosphere during transit

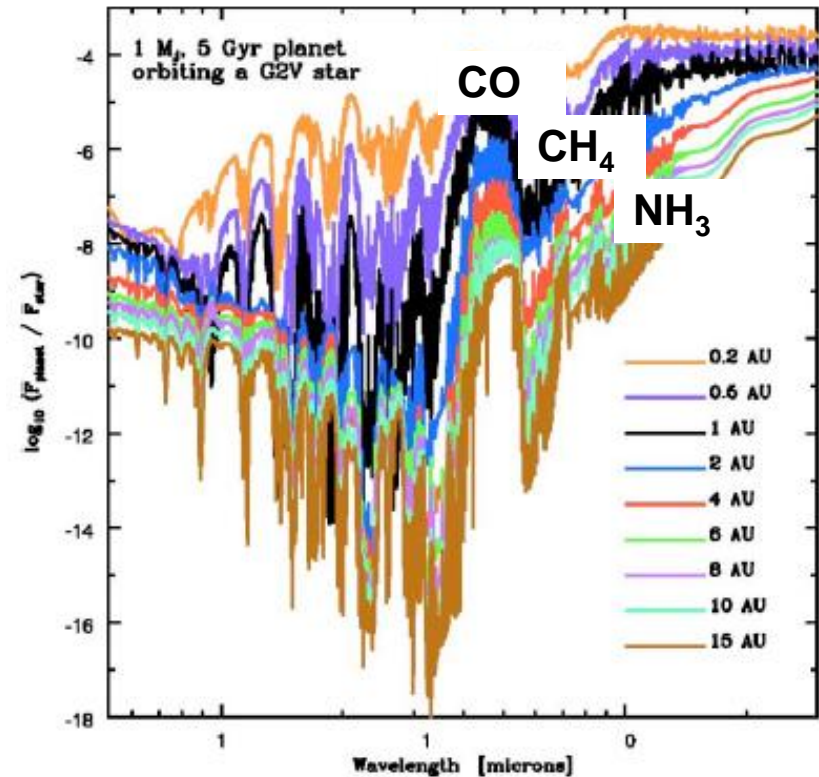
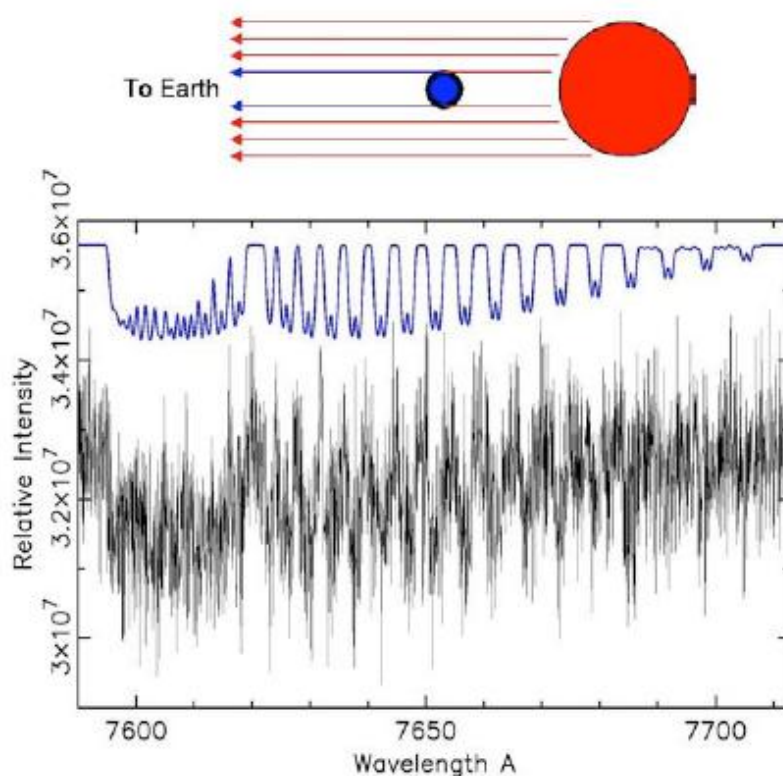
ハビタブルゾーン (生命居住可能領域)

= 液体の水が存在



Planetary atmosphere

- ◆ Absorptions due to molecules in the transiting planetary atmosphere are superimposed on the spectrum of the host star.
- ◆ 3hrs integ. (snr=30000, 6km/s) of TMT/HROS for O₂ of Mstar
- ◆ MIR high. res. Spec. for organic hydrocarbon molecules



◆ **MIR+NIR+Opt Users Migration**

- ◆ Exploding Akari/SPICA User population
 - fostering native MIR youngsters
 - encouraging **redshift immigrants** from Subaru community
 - inviting blueshift immigrants from ALMA community
 - inviting foreigners from Physics/Planetary Science communities
 - (bringing targets, ideas, instruments,)
 - ◆ Emigrants to Subaru/ TMT
 - NIR/Opt science for MIR targets
 - Participation in TMT instrumentation
 - (No channel for cosmological blueshift immigrants)
- 14/49 TMT SWG members are also SPICA SWG members**
- TMT SWS in Victoria (late March- mid April?) : status reports**

◆ SPICA SCIENCE MANIFESTO

SUPPORT FROM FUNDING AGENCIES

◆ SPICA SCIENCE PRIORITY

◆ MIR baseline

◆ Bonus to FIR or NIR

◆ Clarifying science objectives

(differences among ALMA, TMT, SPICA)

◆ International Partnership

◆ Communities tie with ALMA and Subaru/TMT

◆ NAOJ's future roles in Space Astrophysics

discussion at NAOJ Research Program Adv. Comm

◆ Importance of having a coherent future strategy for NAOJ+ISAS .