

Z203a Importance of Dust in the Evolution of Galaxies: Prospect for SPICA

竹内 努 (名古屋大), SPICA サイエンス検討会銀河・BH進化班 (泉 拓磨, 今西 昌俊, 久保 真理子, 播金 優一, 馬場 俊介, 山下 拓時 (NAOJ), 田村陽一 (名古屋大), 梅畑 豪紀 (理研), 鳥羽 儀樹 (京都大), 長峯 健太郎 (大阪大), 橋本 拓也 (早稲田大), 市川 幸平 (東北大), 和田 武彦 (ISAS))

Dust plays various important roles in galaxy evolution. First, dust grains are tiny particles of heavy elements, and they should directly reflect the evolution of galaxies. The dust grain surface works as a catalyst for molecular formation, leading to the first burst of star formation. The stellar emission is strongly attenuated by dust, and re-emitted at mid–far infrared (M–FIR). We constructed a framework of dust evolution based on the chemical evolution (Asano et al. 2013a, b; 2014; Nozawa et al. 2015). Then we extended it to include the infall of baryons (Nagasaki 2020). We also developed a radiative transfer model of galaxy spectrum based on Asano+ framework with a Mega-Grain approximation (Nishida 2020). This provides us with a convenient set of theoretical tools to explore the dust in galaxies at any redshift from various aspects. After the advent of SPICA, we will be able to tackle some fundamental and interesting problems on the dust evolution in galaxies. Since we have an evolutionary radiative transfer model, the astromineralogy of galaxies will be feasible. We will be able to decompose the dust species into several silicates and carbonaceous grains especially from the MIR spectra obtained by SPICA. We should also resolve the so-called “dust budget crisis”, a problem that very young galaxies at extremely high- z have too much dust. The recently proposed scenario of dust formation by AGNs will also be explored by SPICA. We will present the promise of SPICA from the theoretical side.