

Subaru-WFIRST Synergistic Observations

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(* under ISAS Space Science Advisory Committee)

WFIRST-Subaru Synergistic Observations Workshop

December 18-20, 2017

http://www.ir.isas.jaxa.jp/WFIRST_Subaru/index.html



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Important Dates

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First announcement — save the date

WFIRST-Subaru Synergistic Observation Workshop

to be held from December 18 - 20, 2017

at

Large Seminar Room, NAOJ, Mitaka, Tokyo, Japan

Potential Japanese Contribution "Package" for WFIRST

- 1. Subaru-WFIRST Coordinated/Synergistic Observations
- Potential Contribution to Coronagraph Instrument Polarimetry capability
 w/ Polarimetry Compensation Unit
- 3. Ka-band Data Downlink Station in Japan
- 4. Ground-based microlensing observations:
 - MOA dataset for development of the pipeline
 - pre/concurrent observations with a new 1.8m telescope
 (PRIME: a dedicated telescope for microlensing)

Polarimetry Capability for CGI

- Polarimetry Unit (Imaging Polarimetry)
- Development of Polarization Compensator correcting polarization-differential wavefront aberration (PDWA)

Adding accurate polarimetry capability

- important science cases for planets and-disks
- achieving higher contrast

Development of Polarization Compensator

N. Murakami, et al.

Broadband high-contrast polarimetry observations

Problem:

- Instrumental polarization causes polarization-differential wavefront aberration (PDWA)
- DM cannot correct different X- and Y-polarized WFs simultaneously

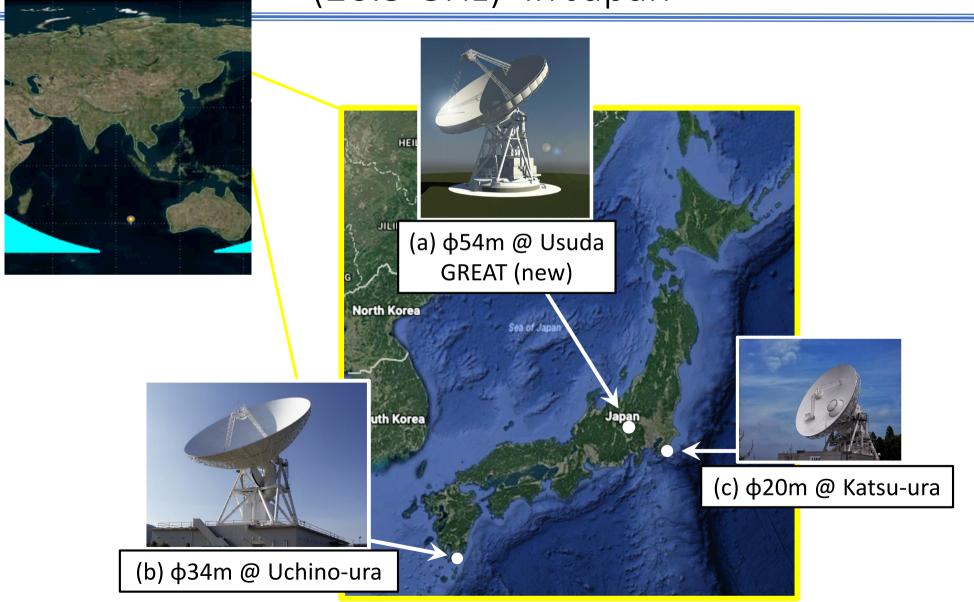
Polarization-compensating system

Birefringent plates: reduce the PDWA to make

the X- and Y-polarized WFs be identical

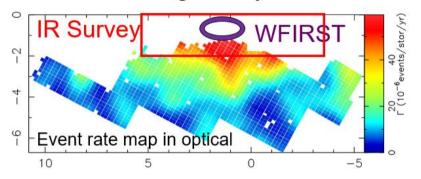
Non-birefringent plates: correct (flatten) the distorted WFs

Ground Station Option for Ka-band (26.5 GHz) in Japan

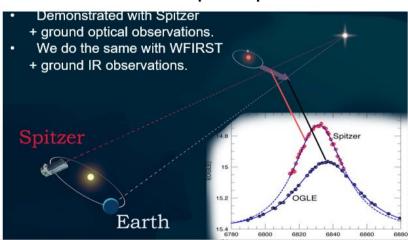


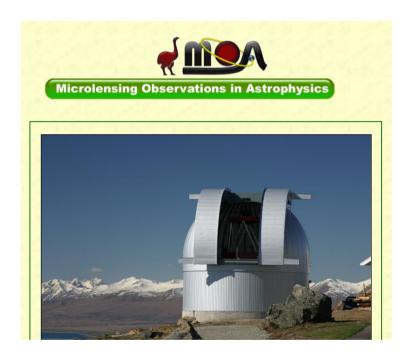
Ground-based microlensing observations: • MOA dataset for development of the pipeline

 Precursor observations for optimizing WFIRST microlensing survey field



2. Concurrent observations with WFIRST for lens mass determination via Ground-Space parallax









Current Status (1)

- ✓ Subaru/WFIRST synergistic observations NAOJ agreed to commit ~100 Subaru nights at around 2025 for the Subaru/WFIRST synergistic observations.
 - NAOJ is ready to send the Letter of Commitment to JAXA, which will be directly cc-ed to NASA
 - Subaru/WFIRST Workshop in Tokyo Dec 18-20, 2017
- ✓ 'MOA-Project' microlensing data sharing
 MOA project agreed to commit sharing the full
 microlensing survey data with the WFIRST tam

NAOJ: National Astronomical Observatory of Japan, which operates Subaru MOA: an international microlensing survey consortium led by Japanese universities





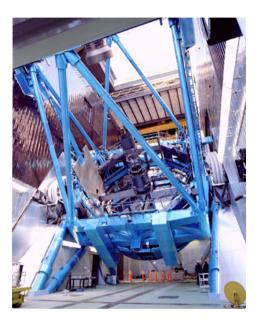
Current Status (2)

- ✓ Ka-band Ground-Station
 - No current Ka-band facility of JAXA
 - JAXA can provide Ka-band data downlink capability (and S-band communications) for WFIRST, by upgrading either Usuda new 54m antenna (2020-) or the Uchino-ura (USC) 34m antenna
 - Discussion with NASA WFIRST team for planning
- Polarimetry imaging capability for the Coronagraph Instrument is under study with the NASA JPL team
- Plan for the new microlensing-dedicated 1.8m telescope is under discussion

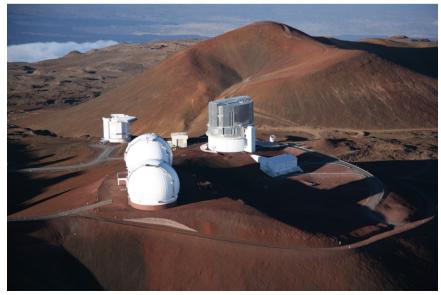


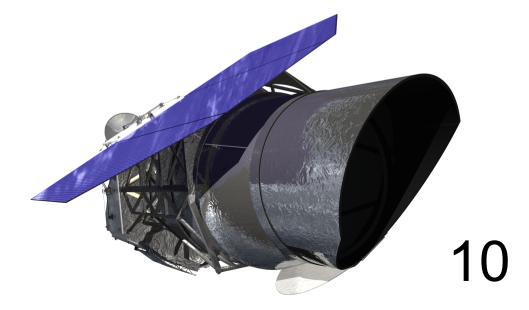
Subaru/WFIRST Synergistic Observations





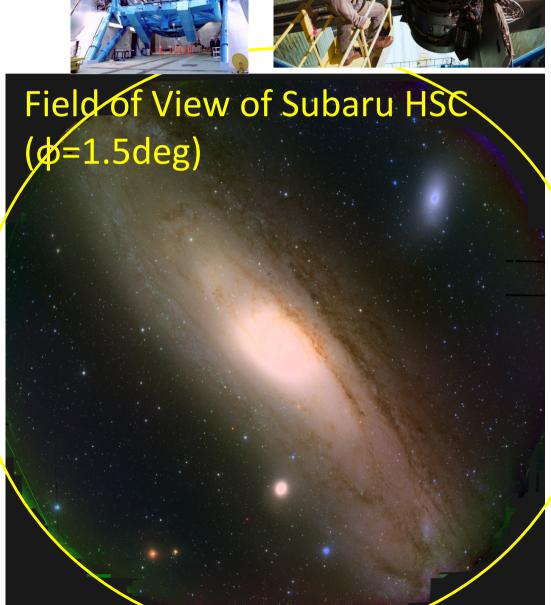
- ➤ Northern Hemisphere (←→ LSST)
- Optical Wavelength (NIR for WFIRST)
- Wide-field Imager (Hyper Supreme Cam; HSC)
- ➤ Wide-field Spectrograph, 2020~ (Prime Focus Spectrograph; PFS)

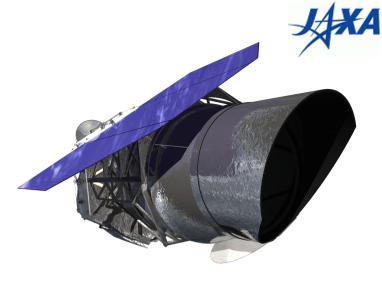














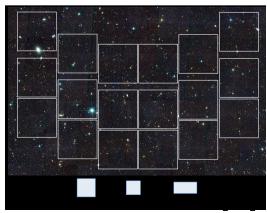
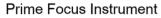
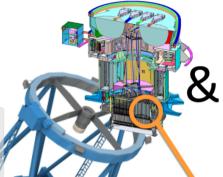


Figure 1-1: Field of view comparison, to scale, of the WFIRST-AFTA wide field instrument with wide field instruments on the Hubble and James Webb Space Telescopes. Each square is a 4x x 4k HgCdTe sensor array. The field of view is 0.22 degrees. The pixels are mapped to 0.11 arcseconds on the sky.







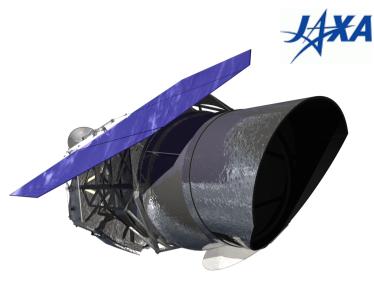






Subaru PFS
Spectrograph
Ф1.3deg
2400 fibers







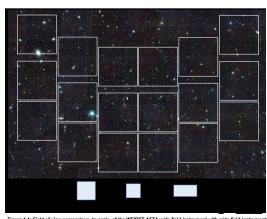


Figure 1-1: Field of view comparison, to scale, of the WFIRST-AFTA wide field instrument with wide field instruments on the Hubble and James Webb Space Telescopes. Each square is a 4x x 4k HgCdTe sensor array. The field of view is 0.28 degrees. The pixels are mapped to 11 arcseconds on the sky.





Subaru/WFIRST Synergistic Observations

Science Synergy of Subaru and WFIRST

Some voices..

- E.g., Subaru PFS deep and systematic spectroscopy can provides the redshift distribution of the galaxies to calibrate WFIRST cosmology analysis.
- E.g., Redshift for Supernova hosts
- E.g., Subaru HSC and PFS can provides optical broad-band and narrow-band imaging and spectroscopic information for the WFIRST "Deep Survey"
- Feeding targets to TMT at north





Subaru Synergistic Observations

- Subaru users show great interest in the synergistic observations with WFIRST
- Subaru Telescope can reserve a certain number of nights (~100 TBD) at ~2025 for the Subaru-WFIRST synergetic program, if it is supported by the Subaru community.
- Letter of Intent from Nobuo Arimoto, Director of Subaru Telecope to ISAS/JAXA Director General Saku Tsuneta, which is CC-ed to Dr. Paul Hertz of NASA, Sept, 2016
- Good support in the GOPIRA symposium (Sept 27,2016)
 GOPIRA=group of optical and infrared astronomers

■Letter of Intent from Nobuo Arimoto, Director of Subaru Telecope to ISAS/JAXA Director General Saku Tsuneta, CC-ed to Dr. Paul Hertz of NASA.

Subaru Advisory committee supported

Subaru Community Discussions
2015.12 Presentation at Subaru Advisory Committee (SAC)
2016.1 Subaru Users' Meeting (w/Jason Rhodes)
2016.2 Presentation at Pasadena Meeting
2016.4 SAC workshop for synergy with Space Mission in 2020's
(mainly, WFIRST, and Euclid + TESS)
2016.5 report for the workshop / SAC news letter 62
2016.6 Presentation ag WFIRST FSWG #2
2016.7 Update at Subaru Advisory Committee (SAC)
2016.9 Supported at the GOPIRA Symposium

2016.10 SAC news letter 67 strong support for the commitment

- Details for the Workshop
- 2016.6 Subaru-WFIRST synergy whitepapers (Japan)
- 2017.6 Discussions in WFIRST FSWG#2
 - → Proposal for the workshop in Japan, in December
- 2017.7 Workshop announced to GOPIRA
- 2017.10 Japanese Discussions #1
- 2017.11 Japanese Discussions #2
- 2017.12 Workshop on Dec18-20

 As for the Japanese contribution to WFIRST, NAOJ, Subaru Telescope, and Subaru and Opt/IR community have agreed to commit for ~100 nights at around 2025 of Subaru Telescope for WFIRST-Subaru Synergistic Observations.

 The purpose of this workshop is therefore to survey the requirements from both WFIRST and Subaru side in Surveys and GO various science fields.

 WS provides the first opportunity for WFIRST and Subaru communities and the representative scientists to discuss and identify the most valuable way to use of the committed nights.

 This is also the first important opportunities for Japanese researchers to learn and consider the science with WFIRST.

Some Important Notes for Subaru-WFIRST time

- When the 100 nights will be scheduled
 It is difficult to assign a large number of nights
 before 2024 (as PFS Strategic Program runs 2019-2024+).
 We may assume 100 nights will be scheduled over a few years,
 possibly over the 4 years, the nominal program of WFIRST
- ➤ A fraction can be used before ~2025?
- Coordination with Subaru Open Useconstraints from the observatory (see Yoshida san's talk)
- ➤ Balance of dark/grey nights
- > Possibility to consider much larger number of nights

Some Important Notes for Subaru-WFIRST time

> Steering

Subaru-WFIRST Time Steering Board should be set

- WFIRST WG rep, Subaru Committee rep, Subaru Telescope rep., JAXA rep
- Project Scientists, SWG/SIT reps.

(discussion on the 3rd day)